International Conference on Instrumentation for Colliding Beam Physics

**INSTR - 2017** February 27 - March 3, 2017 BINP, Novosibirsk, Russia

Physics (INSTR 2017) will be held in the Budker Institute of Nuclear Physics, Russia from February 27 to March 3, 2017. Since 2012 this conference is in a series of Instrumentation conferences held by rotation on Elba (PM or Pisa Meeting on Advanced Detectors), in Vienna (VCI or Vienna Conference on Instrumentation) and Novosibirsk.

#### SCIENTIFIC PROGRAM

The aim of the Conference is to review the status and progress in instrumentation for experiments at colliding beams and related fields. The main topics include:

. Colliders and detector integration

- . Tracking and vertex detectors
- Micropattern gas detectors
- Particle identification
- Calorimetry
- Instrumentation for Astroparticle and Neutrino physics Electronics, Trigger and Data Acquisition

#### e deadline for registration - Janu

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N \*Novosibirsk State University THE REAL SCIENCE



#### The **BELLE** Electromagnetic Calorimeter and its upgrade to BelleII C. Cecchi University of Perugia and INFN-PG



#### Outline:

- BELLE CsI(TI) calorimeter
- SUPERKEKB and Bellell
- Front End electronics upgrade for Bellell
- Commissioning of ECL for phase 2 and 3
- Upgrade of the CsI(TI) to pure CsI
- Summary and Conclusions



#### Electromagnetic calorimeter @B-factories

- Large dynamic range: 20MeV~8GeV
  - 1/3 of B decays have  $\pi^0$ , most of  $\gamma$  ~100MeV.
  - Radiative B decays ( $B \rightarrow K^* \gamma$ , etc.)  $\gamma$  up to 4GeV
  - Bhabha, e+e-  $\rightarrow \gamma \gamma$  calibration, up to 8GeV
- High energy resolution
  - $\sigma_{\rm E}/{\rm E}$  ~ 4% at 100MeV
  - $\pi^0$  Mass resolution ~5MeV
- High position resolution
  - $\sigma_x$ : 5~10mm at the incident point

## BELLE CsI(Tl) Calorimeter



INFN



#### Performances









#### ECL in view of luminosity increase

Radiation damage of crystals after BELLE (1 ab<sup>-1</sup>)



20

0

01/03/2017

**40** 

60

 $\Theta_{id}$ 

ecchi

Expected dose in crystal @ Belle-II FWD ECL: 5-6 Gy/yr x 10 yr of data taking  $\rightarrow$ a factor of 13 higher than in BELLE

In the most loaded part the light output degradation is about 10%

See D.Matvienko's poster



### Background increase: pile-up effect



(E>20 MeV) 6 fake clusters, 3 in barrel 3 in endcaps background



#### Background MC campaign



#### 14.5th campaign maximum dose: 5.52 Gy/yr

15th campaign maximum dose: 6.27 Gy/yr



# ECL pulse-shape calibration



- First calibration with very preliminary start coefficients is done by about 130 iterations and takes about 30 min. The next calibrations take about 10 min.
- Whole calibration time (including data taking and copying raw files to KEKCC) takes about 1.5 days.





### Endcap Installation



Endcap tested before and after installation  $\rightarrow$  all channels working and all DSP shapers

FWD Endcap installation September 2017

Barrel and Backward are being monitored and calibrated (barrel) → stability 10<sup>-3</sup>





### ECL cosmic test

- CDC-ECL cosmic run was performed Feb. 2 2017
- Short runs with total statistic of about 4000 events
- ECL distributions seems to be reasonable





#### Environment monitoring





 $\sigma = 0.004490 \pm 0.000026$ 

-94000

### New ECL reconstruction



01/03/2017

C. Cecchi



## FWD Endcap Upgrade pure CsI

#### Electronics upgrade may not be sufficient for the forward endcap

Crystal	$\operatorname{Csl}(\operatorname{Tl})$	Csl
Density $(g/cm^3)$	4.51	4.51
Melting Point ( $^{\circ}CC$ )	621	621
Radiation Length (cm)	1.86	1.86
Molière Radius (cm)	3.57	3.57
Interaction Length (cm)	39.3	39.3
Refractive $Index^a$	1.79	1.95
Hygroscopicity	$\operatorname{Slight}$	Slight
Luminescence <sup>b</sup> (nm) (at Peak)	560	$\begin{array}{c} 420\\ 310 \end{array}$
Decay Time <sup><math>b</math></sup> (ns)	1220	$\frac{30}{6}$
Light Yield <sup><math>b,c</math></sup>	165	$\begin{array}{c} 3.6 \\ 1.1 \end{array}$
$d(LY)/dT^{b,d}$ (%/°CC)	0.4	-1.4

Because of short scintillation decay time, ~30ns, Pure CsI crystal is almost pileup free. Photo Pentode readout is regarded as a baseline, noise~0.2MeV.

- same density and radiation length allow to reuse Belle mechanical structure
- much lower light yield
  - fast component of emitted light in the near-UV region
- much faster light decay time
  - but slow component is an issue for pile-up



- Use of Pure Csl requires R&D studies on:
  - photodetectors in the near-UV and wavelength shifters
  - radiation hardness of crystals, photodetectors, and wavelength shifters
  - electronics

01/03/2017



## Pure CsI optical properties:

#### transmittance

- No irradiation
- @ 310 nm: L.T. ~ 50%, T.T. ~ 40% for Optomaterials, lower for Amcrys

- Transverse transmittance at different irradiated doses (7 to 104 Gy)
- Maximum variation @ 310 nm
   -7%
- Saturation effects
- Completed and fast recovery for doses < 7 Gy (not shown here)





#### Pure CsI Optical properties: Light Yield



- Longitudinal LY variation (Irradiated)
- before irr. : ~ 6-10%
- after irr: ~ 7-15%
- LY variation before and after irr.: ~1-5%
- No irradiation
- Longitudinal LY variation < 10%</li>



## Photodetector options

Hamamatsu

Requirement: photodetector ENE < pile-up noise ( O(1MeV) )

#### Photodetector options:

	Hamamatsu photonics R11283 photopentode Belle-II baseline	APD R8664-1010	
	PhotoPentodes	Large Area APD (LAAPD)	
C (pF)	10	270	
Gain@V <sub>op</sub>	150-250 [*]	50 [**]	
<u>Q.E.@310</u> nm (%)	25	50	
Area (cm²)	20	10x10	
Comments	<ul> <li>[*] gain reduced by 75% in 1.5T magnetic field</li> <li>1 PP per crystal, no redundancy</li> <li>back plane of the mechanical structure need to be replaced</li> </ul>	<ul> <li><sup>[**]</sup> special production with G=200</li> <li>2 LAAPD per crystal</li> </ul>	



### R&D on LAAPD

8 200

180

160

140

120

100

80

60

40 20

0.05

LAAPD sum, Max Amplitude in full range, run 168

0.1

0.15

hn ha

793

0.1439

0.0516

7938

0.1439

0.04656

178.3 ± 3.4

 $0.01583 \pm 0.00036$ 

0.3

15 - 0.0002

Entrie

Mean RMS  $\chi^2$  / ndf

Prob

Const

Entries

Consta

0.25

Amplitude (V)

Mean

RMS

0.2

Extensive R&D has been done on Pure CsI + LAAPD

#### They meet the experiment requirement (ENE O(1MeV))

σ[mV]	Signal [mV]	S/N	ENE[MeV]
2.80±0.17	108.6±0.6	38.8±2.4	0.77±0.05



- gamma: 250 Gy = 10 y data taking x ~10 safety factor
- Neutron fluence:  $10^{12}$  n/cm<sup>2</sup> = 10 y data taking x 5 safety factor





 Novel wavelength (WLS) plates containing nanostructured organosilicon luminophores provides essential increase in light output



• Results:



Enhancement on signal of a factor of about 3 (n.b. test per <sup>01/02/2017</sup>formed with G=50 LAAPD)





Radiation hardness tests on NOL9 WLS: no irradiation effects on excitation/emission peaks up to <sup>C. Cecchi</sup> 105 Gy



#### Pile-up study: cosmic+source

- Signal from cosmic rays superimposed to 1.33 and 1.17 MeV g produced by <sup>60</sup>Co
  - 1.77±0.04 hits/µsec from source
- From the difference in quadrature of the meas. with and without source, pile-up effect can be estimated: CsI(TI) shows twice the ENE of pure CsI
- [nb: wrt to pag 12 (ENE=0.77MeV), different shaping times and signal window used, optimization of such parameters underway ]





### Conclusions

• Belle CsI(TI) calorimeter successfully operated for 10 years to bring various physics output

• SuperKEKB is aiming x40 luminosity w.r.t. KEKB.

- Beam background control is a challenge for the detector and its performances

Use all the existent CsI(TI) with waveform sampling readout electronics. Full simulation studies for CsI(TI) deterioration  $\rightarrow$  new reconstruction algorithm

• PureCsl upgrade

-Extensive R&D on crystals and photodetectors shows that the upgrade represents an improvement in dealing with beam background and pile-up - a real estimation of the background is necessary before the decision is taken  $\rightarrow$  after the first run of Bellell for the background measurement