

# Combined Xe – CsI calorimeter of CMD3 detector

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on behalf of CMD3 collaboration

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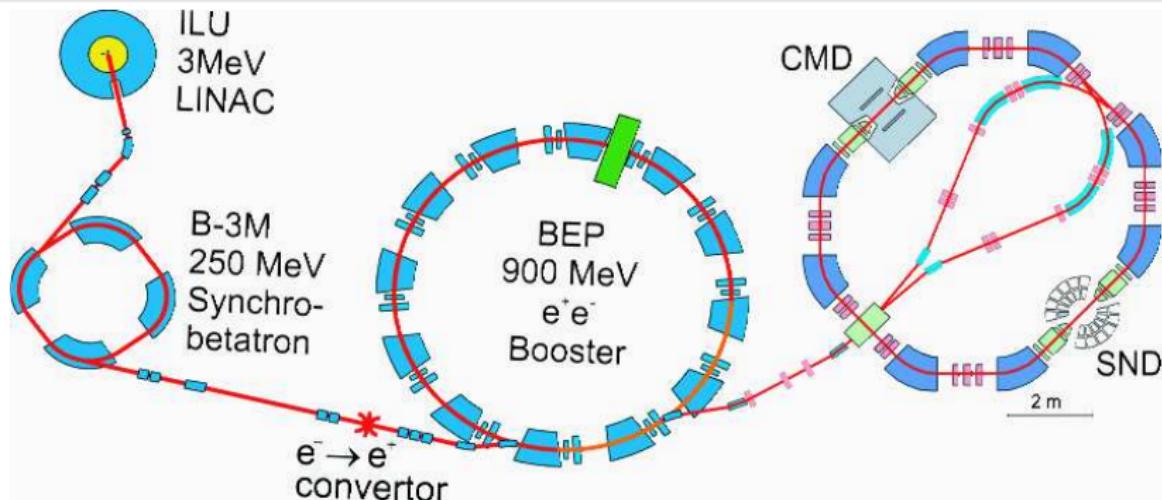
INSTR2014



# Outline

- CMD-3 detector
- LXe calorimeter
- CsI calorimeter
- Cosmic calibration
- Bhabha calibration
- Clusters reconstruction
- Photon energy and coordinate correction
- Conclusion

# VEPP-2000

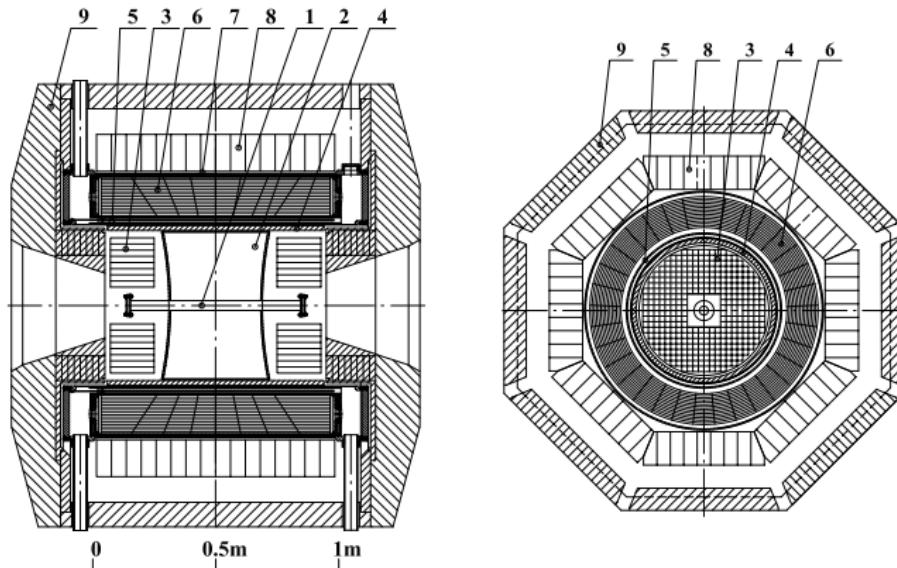


## VEPP-2000

- $e^+e^-$  collider
- $E_{c.m.s.} = 0.3 \div 2 \text{ GeV}$
- SND and CMD3 detectors are mounted at interaction points
- $L_{\text{project}} = 10^{32} \text{ cm}^{-2} \text{s}^{-1}$ ,  $L_{\text{exp.max}} = 2 \cdot 10^{31} \text{ cm}^{-2} \text{s}^{-1}$
- 2010-2013 year of operation:  $L_{\text{int}}(\text{CMD3}) \approx 60 \text{ pb}^{-1}$



# CMD3 detector



1 – beam pipe, 2 – drift chamber(DC), 3 – BGO endcap calorimeter, 4 – Z-chamber,  
5 – superconducting solenoid, 6 – Liquid Xe(LXe) barrel calorimeter, 7 – TOF, 8 – CsI barrel  
calorimeter, 9 – Yoke

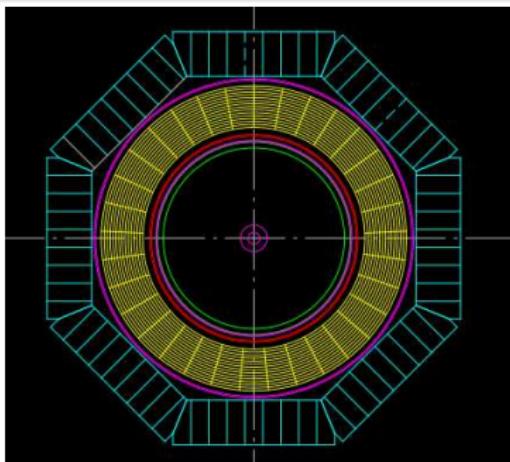
- Magnetic spectrometer,  $B = 13\text{kGs}$
- Thin superconducting solenoid is mounted in the same vessel with LXe calorimeter.  
Passive material in front of barrel calorimeter is  $0.35X_0$  ( $6.27\text{g/cm}^2$ )
- 3 calorimeter systems: barrel LXe and CsI calorimeters and endcap BGO calorimeter



# Combined barrel calorimeter

## Tasks

- Measurement of photons energy in wide range from 10MeV upto 1GeV
- Measurement of photons coordinates
- Form a signal for neutral trigger



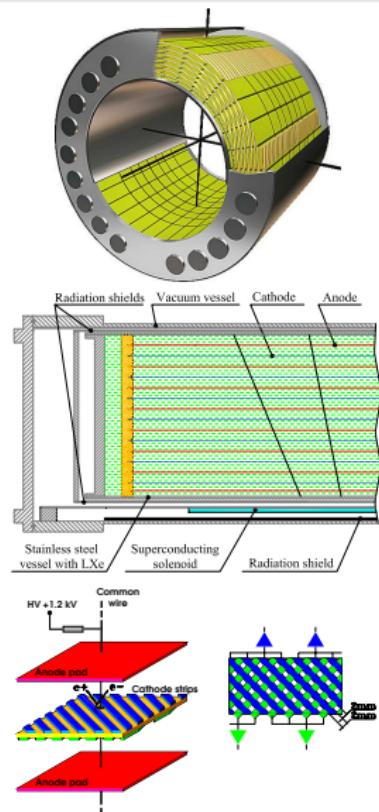
## Advantages

- LXe strips structure provides high spatial resolution, allows measurement of specific energy loss and shower profile
- Full thickness of barrel calorimeter is  $13.5 X_0$  and provides energy resolution of 8% for 200MeV and 4% for 1GeV energy deposition

## Disadvantages

- Passive material between LXe and CsI is  $0.25 X_0$

# LXe calorimeter



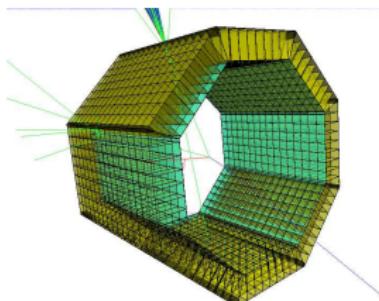
## Calorimeter structure

- 14 cylindrical ionization chambers devided by 15 electrodes with 10.2mm gap
- 8 anodes and 7 cathodes
- Anodes are devided into 264 towers provide energy measurement:
  - 8 towers along Z axis
  - 33 towers along azimuth angle
- Cathodes are devided into 2112 strips to measure coordinate
- Strips on opposite sides of cylinder are mutually perpendicular
- Thickness –  $5.4X_0$

LXe calorimeter is described in details in K.Mikhailov poster

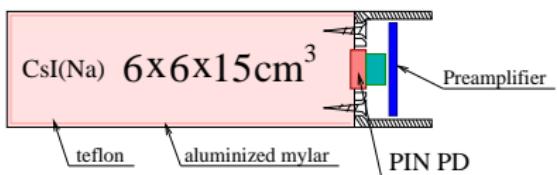


# CsI calorimeter

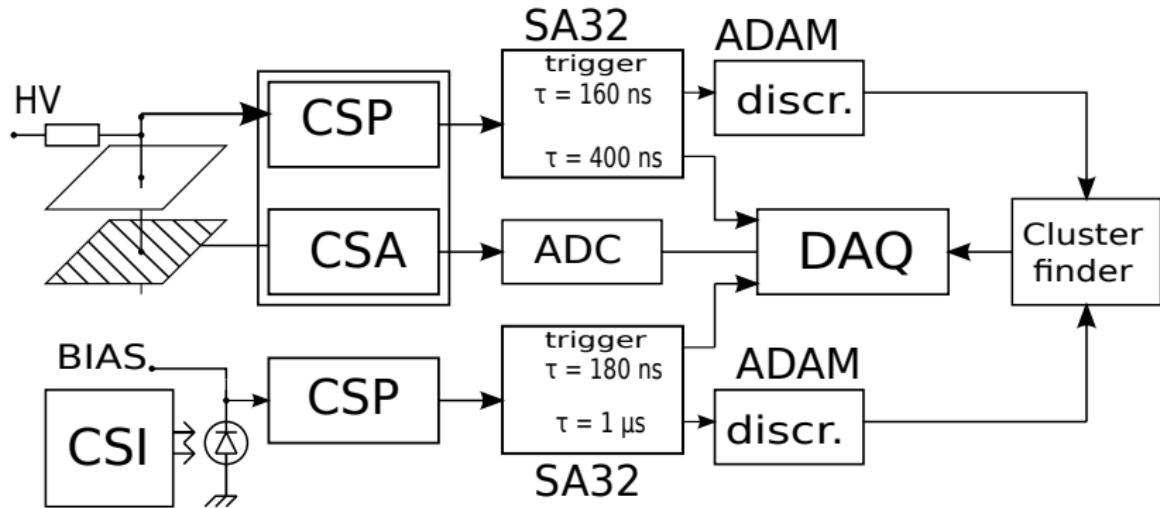


## Calorimeter structure

- 8 octants  $\times$  9 modules  $\times$  16 counters  
**1152 counters in total**
- Counters are made of CsI(Tl) or CsI(Na) crystals
- Scintillation light is read out by Hamamatsu PIN PD mounted on the surface of the crystal
- PIN PD sensitive area is  $10 \times 20\text{cm}^2$
- Active material thickness –  $8.1X_0$



# Calorimeter electronics



# Calibration of energy deposition

## Task

$$E_i = k_i(A_i - p_i)$$

$k_i$  is to be determined

## Calibration steps

- Calibration with pulse generator:
  - $p_i$  determination
  - electronics stability monitoring
- Calibration with cosmic particles of LXe and CsI calorimeters
- Final LXe calibration with events of  $e^+e^-$  elastic scattering (Bhabha calibration)

# Cosmic calibration

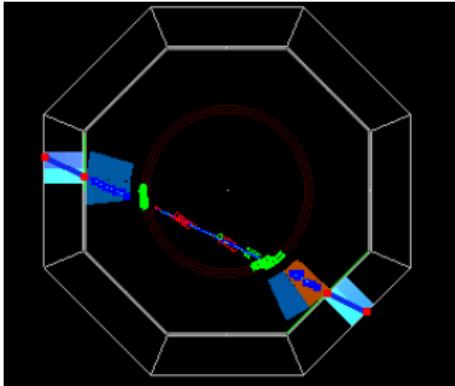
## Method

- LXe track crosses only one LXe tower or CsI crystal
- $3 \times 3$  matrix with the center in the crossed tower(crystal)
- $E_{\text{center}} > 20 \text{ MeV}$
- $(E_{3 \times 3} - E_{\text{center}}) < 10 \text{ MeV}$

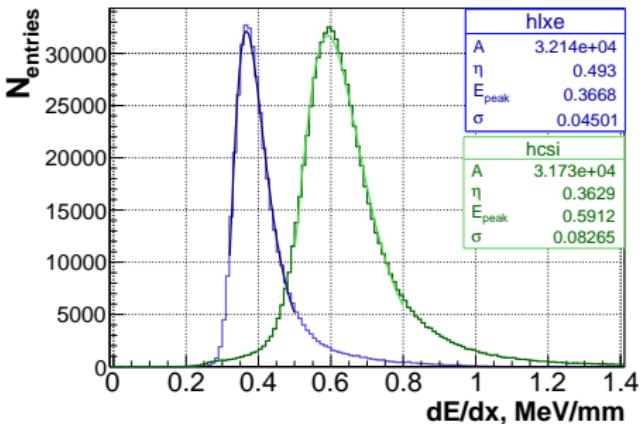
$$\epsilon = E_{3 \times 3} / L$$

$$f(\epsilon) = A e^{\frac{-\ln^2(1+\eta(\epsilon-E_p)/\sigma)}{2\sigma_0}}$$

$$k_i = E_p^{\text{MC}} / E_p^{\text{exp}}$$



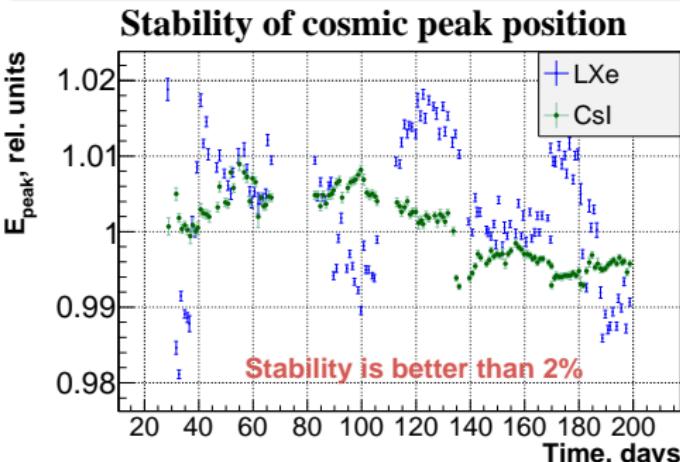
LXe and CsI cosmic spectra (exp)



# Cosmic calibration

- Cosmic events from standard experimental runs that pass the trigger are used. Cosmic events rate  $\nu \approx 70\text{Hz}$
- Calibration for each channel with statistical error of about 1.5% for LXe and 3% for CsI. Data taking takes about 2 weeks.
- Day-by-day calibration – normalization coefficient for 8 LXe towers(same azimuth angle) or for whole CsI calorimeter is measured.

Before day-by-day calibration:

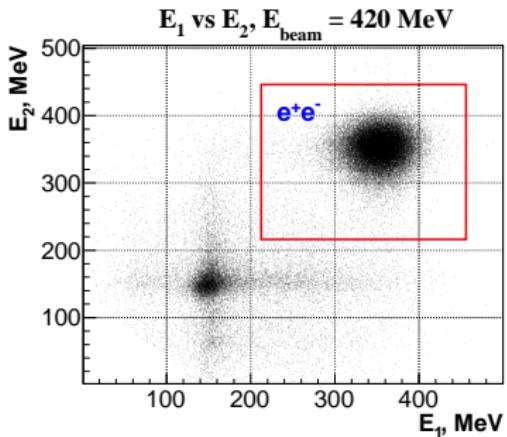


After day-by-day calibration coefficients are applied the stability is better than 0.5%

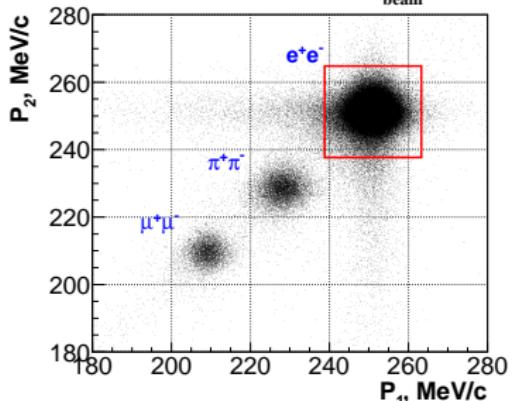
# Bhabha calibration

## Events selection

- 2 tracks in DC
- $|\theta_1 + \theta_2 - \pi| < 0.05$
- $|\phi_1 - \phi_2| < 0.05$
- $|p - E_{\text{beam}}| < 3 \cdot \sigma_p$
- $E_{\text{beam}}/2 < E_{1,2} < 1.05 \cdot E_{\text{beam}}$



Momenta of collinear particles,  $E_{\text{beam}} = 250 \text{ MeV}$



# Bhabha calibration

## LXe coefficients

$$\chi^2 = \sum_{n=1}^N \frac{(E_{mc}^j(\theta, \phi) - \sum E_i^n k_i - E_{CsI}^n)^2}{\sigma^2(\theta)}$$

$$\chi^2 \rightarrow \min$$

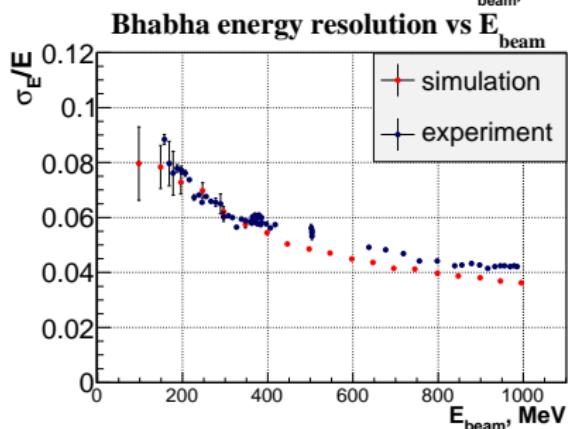
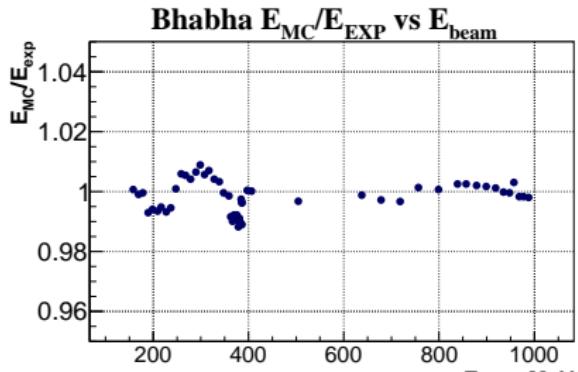
$$\frac{\partial \chi^2}{\partial k_i} = 0 \Rightarrow \sum_j k_j Q_{ij} = R_i,$$

$$Q_{ij} = \sum_{n=1}^N \frac{E_i^n E_j^n}{\sigma^2}, R_i = \sum_{n=1}^N \frac{E_i^n E_{mc}^n}{\sigma^2}$$

$$k_i = \sum_j R_i (Q^{-1})_{ij}$$

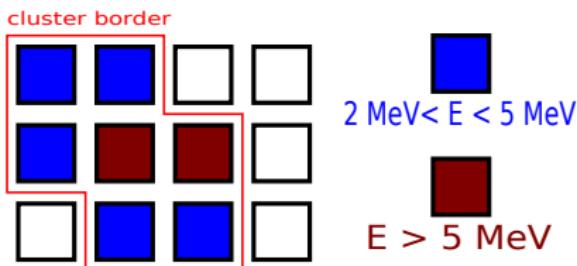
n - event number, i, j - LXe channel indexes

After the calibration is done:



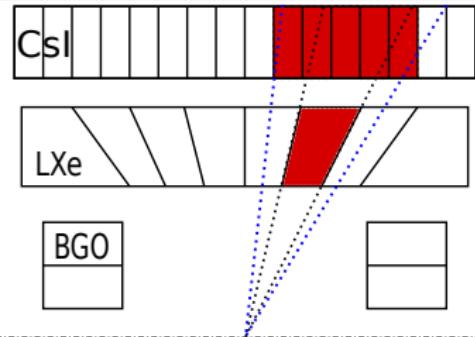
# Cluster reconstruction

Because of the combined structure of calorimeter the clusters in different subsystems have to be connected by some algorithm



## Algorithm(both subsystems)

- For elements with  $E > 5$  MeV, neighbour elements are added to cluster if its energy  $> 2$  MeV
- Elements decided as 'neighbours' if they have common sides



## LXe-CsI connection

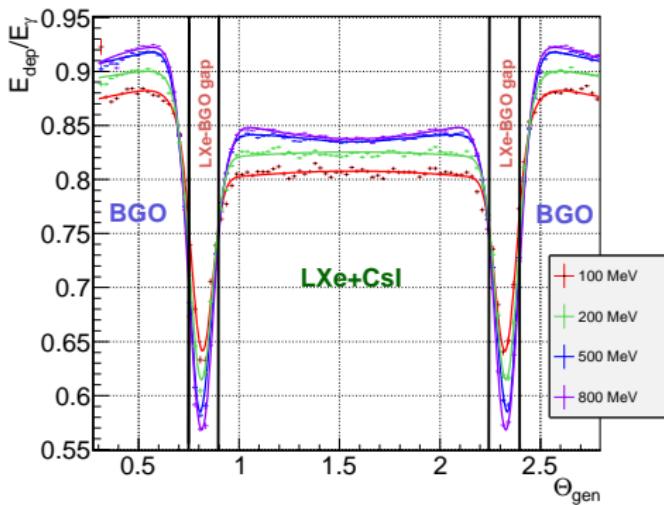
- CsI crystal is 'neighbour' to LXe tower if its center is inside the tower solid angle
- If CsI crystal neighbour to tower is hit, than it is connected to the corresponding LXe cluster



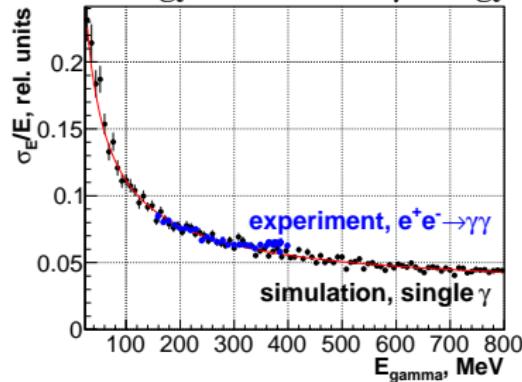
# Photons energy corrections

## $\gamma$ energy

- $E_{\text{gamma}} = f(E_{\text{dep}}, \Theta, \phi)$
- $f(E_{\text{dep}}, \Theta, \phi)$  is determined from simulation of single photons energy deposition



## Energy resolution vs $\gamma$ energy



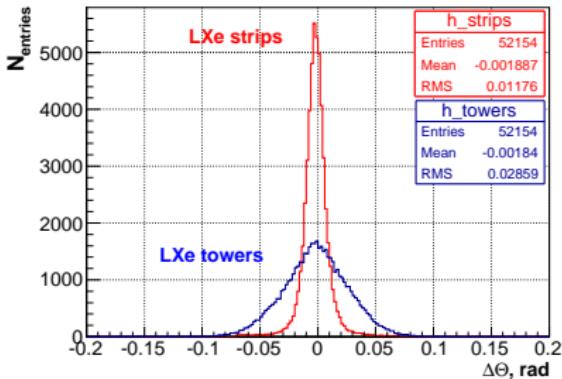
- Energy deposition  $E_{\text{dep}}$  is tabulated for 50 points of photon energy and 200 points by polar angle  $\theta$
- For measured energy deposition  $E_{\text{dep}}$  the photon energy is determined by linear approximation

# Photons coordinate correction

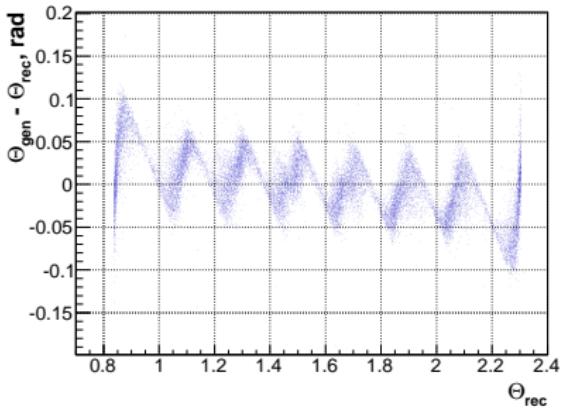
## $\gamma$ angles

- If the gamma conversion point is measured using LXe strips - no correction is needed ( $\sim 95\%$  of photons)
- If there is no strips data:  
$$x_i = \sum_{n=1}^N x_i^n E^n$$
, N - number of LXe and CsI channels in cluster
- Correction function for gamma angles w/o strip data is determined from simulation

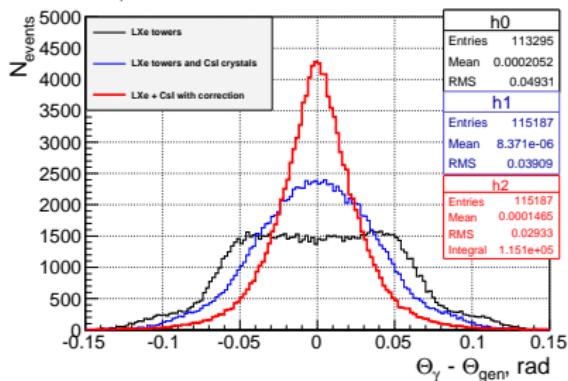
$\Theta_{DC} - \Theta_{LXe}$  distribution (Bhabha)



No strip data for a cluster:

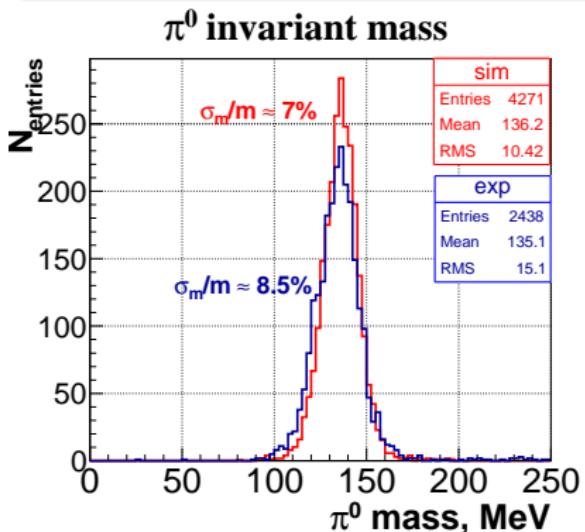


$\Theta_\gamma$  angle resolution (w/o conversion point in LXe)

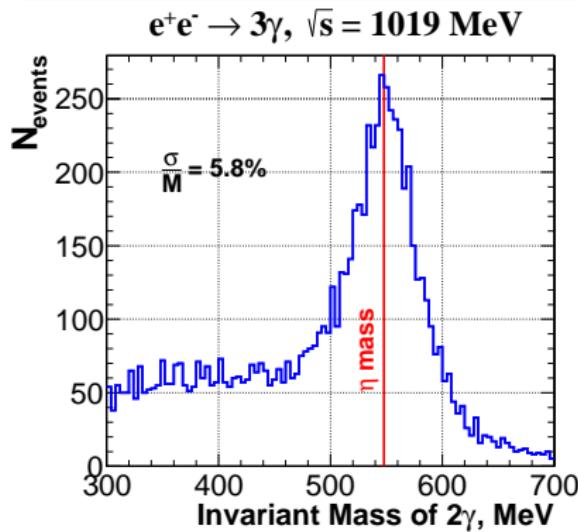


# $\pi^0$ and $\eta$ mass resolution

- $e^+e^- \rightarrow \pi^+\pi^-\pi^0, \pi^0 \rightarrow \gamma\gamma$
- $E_{cm} = M_\phi$



- $e^+e^- \rightarrow \eta\gamma, \eta \rightarrow \gamma\gamma$
- $E_{cm} = M_\phi$

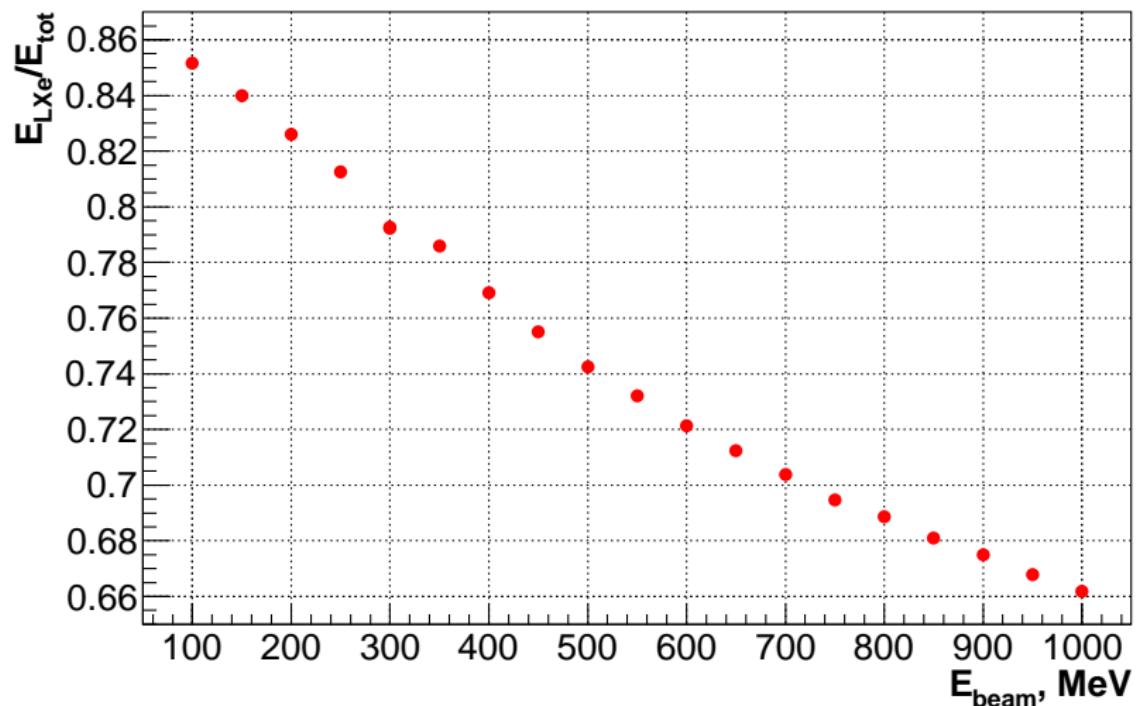


# Conclusion

- Combined barrel calorimeter is working and shows good performance
- Energy calibration procedure of LXe and CsI calorimeters is developed
- Calibration accuracy is about 1% and 3% for LXe and CsI correspondingly
- Energy and coordinate correction functions for photons are determined
- The  $\pi^0$  mass resolution is measured to be 8.5%

# $E_{\text{LXe}}/E_{\text{total}}$ ratio

$E_{\text{LXe}}/E_{\text{tot}}$  vs  $E_{\text{beam}}$  (simulation)



# Photons energy and coordinates resolution

## Energy resolution vs $\gamma$ energy

