Investigation of energy spectrum and chemical composition of primary cosmic rays in 1-1000 PeV energy range with a drone-borne installation

Dmitry V. Chernov¹, Elena A. Bonvech², Miroslav Finger¹,², Michael Finger¹,², Vladimir I. Gal'kin¹,², Dmitry A. Podgudrakov¹,²

1) St Petersburg Institute of Nuclear Physics, M.V. Lomonosov Moscow State University, 122, Leninskii pr., GSP-3, Moscow, 119991, Russian Federation
2) Department of Physics, M.V. Lomonosov Moscow State University, 122, Leninskii pr., GSP-1, Moscow, 119991, Russian Federation
3) Saint Petersburg State University, 199034, Saint Petersburg, Russian Federation
4) Joint Institute for Nuclear Research, Joliot-Curie, 6, Dubna, Moscow region, 141980, Russian Federation

E-mail: chr@dec1.sinp.msu.ru

Abstract
This work is dedicated to the development of a project aimed at the implementation of a relatively new method of studying the PCR – the registration of optical Vavilov-Cherenkov radiation, often called “Cherenkov light”, from EAS (EAS CL), reflected from the snow surface. The objective of the project is to create an installation for the study of the cosmic ray mass composition in the energy range of 1-1000 PeV by detecting the reflected EAS CL. Silicon photomultipliers are planned to be used in the detector of the installation, and an unmanned aerial vehicle (UAV, drone) will be used to lift the measuring equipment over the snow-covered surface.

Introduction
The 1-1000 PeV energy range is transitional from galactic to extragalactic cosmic rays. More than 50 years ago a change in the slope of the energy spectrum of primary cosmic rays (PCR) was detected at around 3 PeV. But until now new features in the structure of the spectrum are being discovered. In this regard it is interesting to understand the cause of the irregularities. The main reason, most likely, are changes in the mass composition of the PCR. Presently used methods allow to estimate either the average mass of PCR particles or to divide them into “light” and “heavy” groups. Basically, estimates are made on the methods allow to estimate either the average mass of PCR particles or to divide them into “light” and “heavy” groups. Basically, estimates are made on the reconstructed depth of development maximum of the extensive air showers (EAS), using the modelling of the shower development in the atmosphere. Since the methodological uncertainties of the reconstructed parameters mass groups can be identified only by processing a large amount of experimental data.

Examples of manufactured drones: the DS900 hexacopter for transporting loads up to 10 kg and a flight time up to 5 hours on a gasoline generator (right).

Previous work
In the period from 2008 to 2013, a series of measurements of reflected Cherenkov light was carried out using the SPHERE-2 [1,2,3] balloon installation. Measurements were made on the snow-covered ice of Lake Baikal.

Advantages of the method
- Provides a significant area of CL registration using a compact device;
- Accurate estimation of PCR energy in an individual event in comparison with other methods;
- The field of view of the individual sensitive elements of the device covers a significant part of the surveyed area, which allows observation of the CL from near the shower axis, usually inaccessible to ground-based CL detector arrays. This circumstance significantly increases the accuracy of the primary particle type estimation;
- Allows measurement of the same PCR energy range with different resolution (distance between the centres of the fields of view of neighbouring sensing elements) using variation of the detector elevation, which allows you to control the magnitude of systematic errors.

Detector
It is planned to design a compact detector that will have the following characteristics:
- Sensitivity of the detector (aperture window) up to 0.1 m²;
- Mirror diameter up to 80 cm;
- Optical system viewing angle 20 - 25 degrees;
- Number of mosaics 49 or more SIPM;
- The mass of the detector less 10 kg;
- The flight height of the detector up to 500 m;
- Expected number of events EAS (with E_0=1-1000 PeV) up to 10,000 for season.

The main element of the new installation will be a segment of seven SIPM Micro FC-60035 SPHERE-2 detectors. Tests of a matrix of seven such segments (49 SIPM) was successfully completed. Each segment was equipped with seven preamplifiers and a temperature sensor to account for the effects of thermal emission. Each SIPM was equipped with a light collector CA1929 and a silicon-on-insulator (SOI) characteristic ±24 degrees at 50% effectiveness. In this project, it is planned to modify and adapt the SIPM segment for use in a ultra-wide angle optical system.

The mirror for the detector will be made on the basis of composite materials. The mirror base is made of aluminium. This design has sufficient rigidity and low weight.

The detector will use the Schmidt optical system. In this system, the central part of the mirror is not used since it is in the shadow of the photodetector. A hole in the centre of the mirror with a wide-angle lens in it with an aperture of ~100 cm² will allow registration of the direct CL. Calculators show that for EAS from PCR ~1PeV the CL photons density is ~100 photons per cm² at a distance of 100 m from the shower axis. Taking into account the SiPM quantum efficiency and losses of optical elements experimental results ~100 photoelectrons. The estimation of the primary particle mass can use the information on the intensity of the direct CL in addition to the data on the reflected CL.

It is assumed that the EAS from the primary proton should form a light spot different of Fe nuclei at the same primary energy and depth of EAS.

To control the density and transparency of the atmosphere an auxiliary small UAV can be used with pressure, temperature, humidity sensors and laser lidar. The lidar will be used to control the reflection from the snow. Laser control of the atmosphere and reflection from the snow will improve the accuracy of measuring the EAS CL.

The properties of the snow surface play an important role when using the method of reflected CL registration. The results of the snow optical properties studies have been repeatedly published by several groups. The simulation results show that in the wavelength range from 300 to 600 nm, the relative reflectance for pure snow is stable within 3% for the zenith angles of the light incidence from 0 to 80 degrees. From the above mentioned results and the known CL spectral characteristics it can be concluded that the snow surface reflects the CL with minor spectral distortions in the range of zenith angles up to 80 degrees and can be used as a screen for the CL registration.

References