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# Low-Temperature Thermal Calorimeters for $0\nu\beta\beta$ experiments

Yong-Hamb Kim (김용함 金容菡)

CENTER FOR  
**UNDERGROUND** PHYSICS

**ibS** 기초과학연구원  
Institute for Basic Science

# Outline

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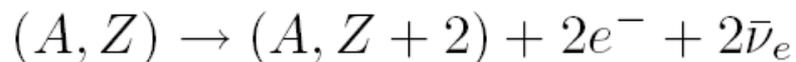
1.  $0\nu\beta\beta$  physics
2. Low-temperature Thermal Calorimeters  
for  $0\nu\beta\beta$  experiments
3. The AMoRE project

# $0\nu\beta\beta$ and $\nu$ (brief intro.)

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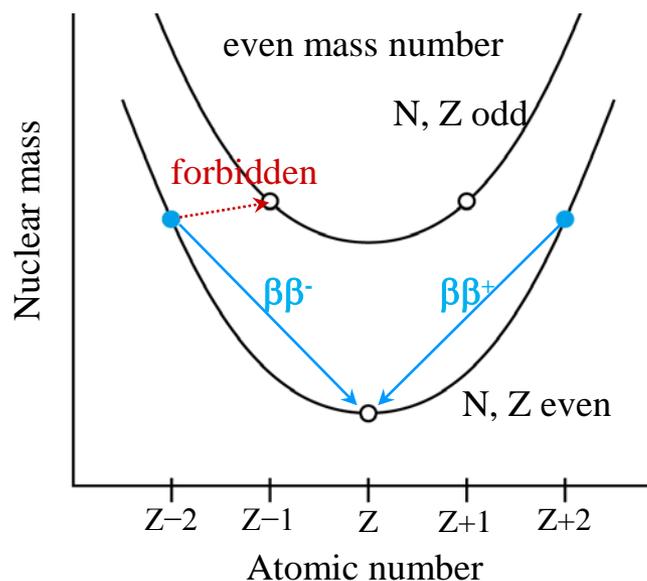
- $0\nu\beta\beta$  decay can only happen if neutrinos are massive Majorana particles (own anti-particles).
  - ✓ fundamental understanding of particle physics
  - ✓  $0\nu\beta\beta$  search is the only practical technique to answer.
- The  $0\nu\beta\beta$  decay rate ( $T^{0\nu}$ ) is closely related to the mass of neutrinos.
  - ✓ Most sensitive measurement method (if Majorana particle)
- The  $0\nu\beta\beta$  decay can only happen if Lepton number conservation is violated.
  - ✓ New physics beyond the standard model

# Double beta decay



35  $0\nu\beta\beta$  nuclei are found.

- 2nd order weak process
- $\beta\beta(2\nu)$  decay is detectable if 1<sup>st</sup> order  $\beta$  decay is not allowed.



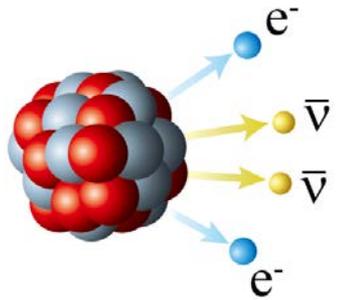
$\beta\beta$ -decay nuclei with $Q > 2$ MeV	Q (MeV)	Abund. (%)
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.271	0.187
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.040	7.8
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Ru}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.034	9.7
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.802	7.5
$^{124}\text{Sn} \rightarrow ^{124}\text{Ge}$	2.228	5.8
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.528	34.2
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.479	8.9
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6

# Double beta decay w. & wo. $\nu$ emission

## 2 $\nu$ mode

- A conventional
- 2nd order weak process in NP

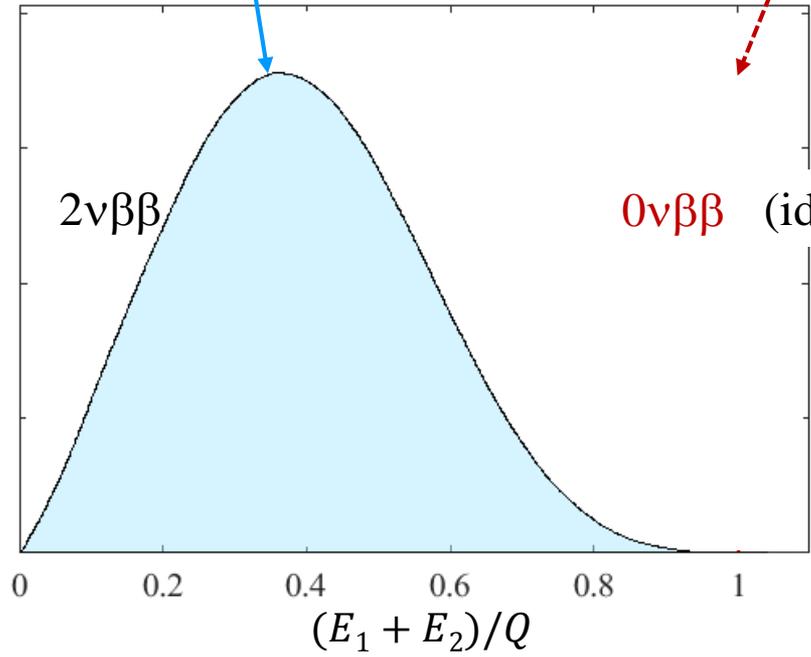
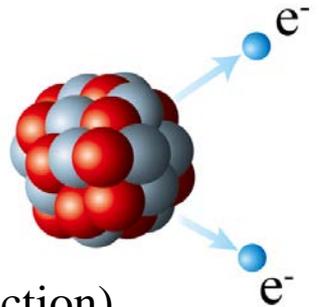
$$(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e$$



## 0 $\nu$ mode

- A hypothetical process only if  $m_\nu \neq 0, \bar{\nu} = \nu, |\Delta L| = 2$

$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$



# Some history about $\beta\beta$ decay



M. Goeppert-Mayer, Phys. Rev. 48 (1935) 512

- ✓ The study of nuclear structure expected that the 2 neutrino mode would have half lives in excess of  $10^{20}$  years
- ✓ First observed directly in 1987.
- ✓ Background:  $_{1/2}(\text{U, Th}) : 10^{10}$  y
- ✓  $_{1/2}(2\nu\beta\beta) : 10^{19\sim 24}$  years



E. Majorana, NuovoCimento14 (1937) 171

G. Racah, NuovoCimento14 (1937) 322

- ✓ The possibility of neutrinos-less decay was discussed in 1937.

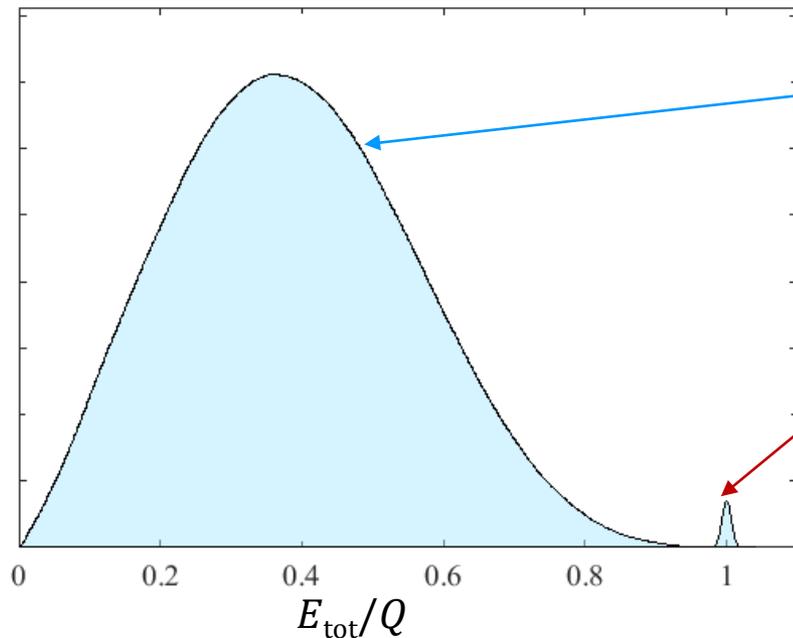


W.H. Furry, Phys, Rev. 56 (1939) 1184

- ✓ Specific discussion on  $\beta\beta$  on Majorana theory ( $0\nu\beta\beta$ )

- ✓ Now, we want to look for a process with  $_{1/2}(0\nu\beta\beta) : \sim 10^{26\sim 28}$  years

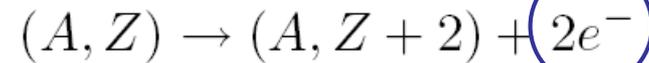
# Search for $0\nu\beta\beta$



**Double Beta Decay with two neutrinos**



**Double Beta Decay with no neutrino**



## $0\nu\beta\beta$ discovery answers

- Majorana ( $\nu = \bar{\nu}$ ) particles not Dirac ( $\nu \neq \bar{\nu}$ )
- Mass scale of neutrinos ( $1/_{1/2}^{0\nu} \propto \frac{2}{\nu}$ )
- Lepton number violation

# $0\nu\beta\beta$ decay rate

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$$\Gamma_{0\nu} = \frac{1}{2} \left| \langle 0\nu\beta\beta | \mathcal{H} | 0\nu\beta\beta \rangle \right|^2$$

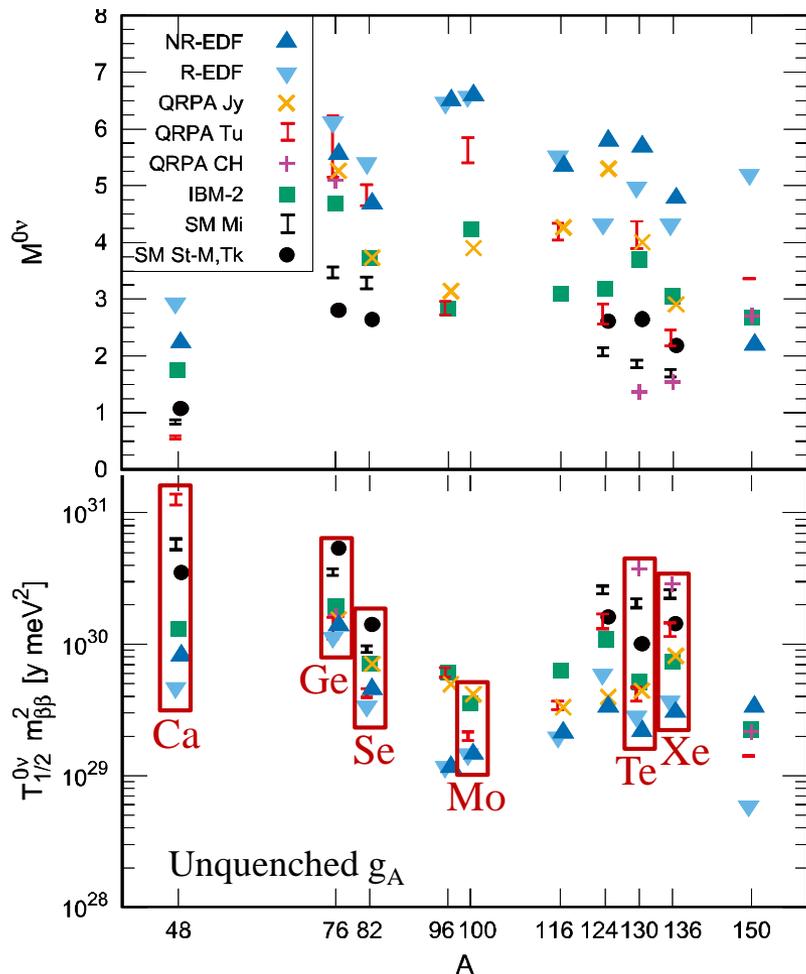
<standard process>

- ✓  $G_{0\nu}$  : Phase space factor: Calculable  
Atomic physics
- ✓  $|\mathbf{M}_{0\nu}|$  : Nuclear matrix element: Uncertain by 2~3 times,  
Nuclear physics
- ✓  $m_{\beta\beta}$  : Effective neutrino mass: Interesting  
Particle physics

# Model dependent NME ( $M_{0\nu}$ )

Engel & Menéndez

Rep. Prog. Phys. **80** (2017) 046301



$$1/T_{1/2}^{0\nu} = G_{0\nu} |M_{0\nu}|^2 m_{\beta\beta}^2$$

✓ Model dependence: 2-3 times spread for each nucleus.

$$1/T_{1/2}^{0\nu} = G_{0\nu} |M_{0\nu}|^2 m_{\beta\beta}^2$$

✓ No significant isotope preference when consider detector mass.

# Another uncertainty in the standard process

$g_A$ : Effective axial-vector coupling constant incorporated in NME

$$1/T_{1/2}^{0\nu} = G_{0\nu} |M_{0\nu}|^2 m_{\beta\beta}^2 \quad \rightarrow \quad \frac{4}{A} G_{0\nu} |M_{0\nu}|^2 m_{\beta\beta}^2$$

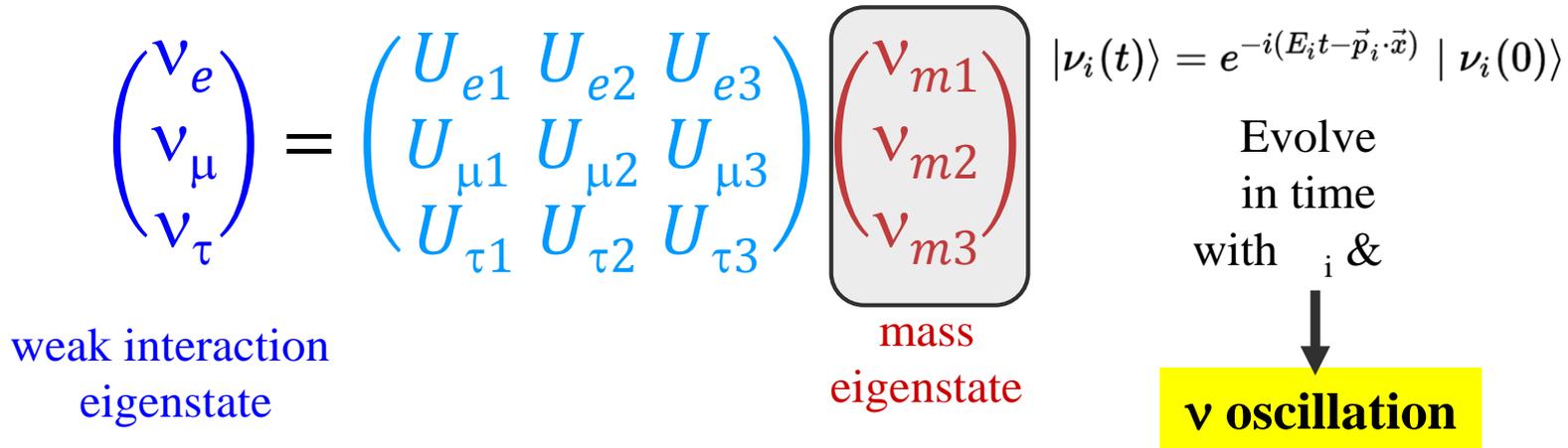
Simple-minded

- ✓  $A = 1.27$  (no quenching for free nucleon)  
= 1 (quark)
- ✓  $A$  is quenched in  $\beta$ ,  $2\nu\beta\beta$ , and possibly  $0\nu\beta\beta$ .  
 $g_A^{\text{eff}}$  for  $0\nu\beta\beta$ : 0.4~0.6

## Effective value of $g_A^{\text{eff}}$ on $0\nu\beta\beta$

- ✓ Theory predictions:  $g_A^{\text{eff}} = 0.6 \sim 0.8$ , <Suhonen PRC 96 (2018) 05550>  
→ Even with small  $g_A^{\text{eff}}$  values,  $0\nu\beta\beta$  decay rates are reduced only by 2-6 times.
- ✓ Some theoretical predictions exist quenching matters with the energy scale  
→ No or less quenching might be needed on  $0\nu\beta\beta$ . <Dolinski et al arXiv:1906.02723>
- ✓ Further theoretical and experimental studies are needed.

# Neutrino mixing, mass, and $0\nu\beta\beta$



## Effective $\beta\beta$ mass

$$\langle m_{\beta\beta} \rangle = \sqrt{\sum_{i=1}^3 m_i^2}$$

virtual  $\nu$  exchange

## Other measurables in $\nu$ mass

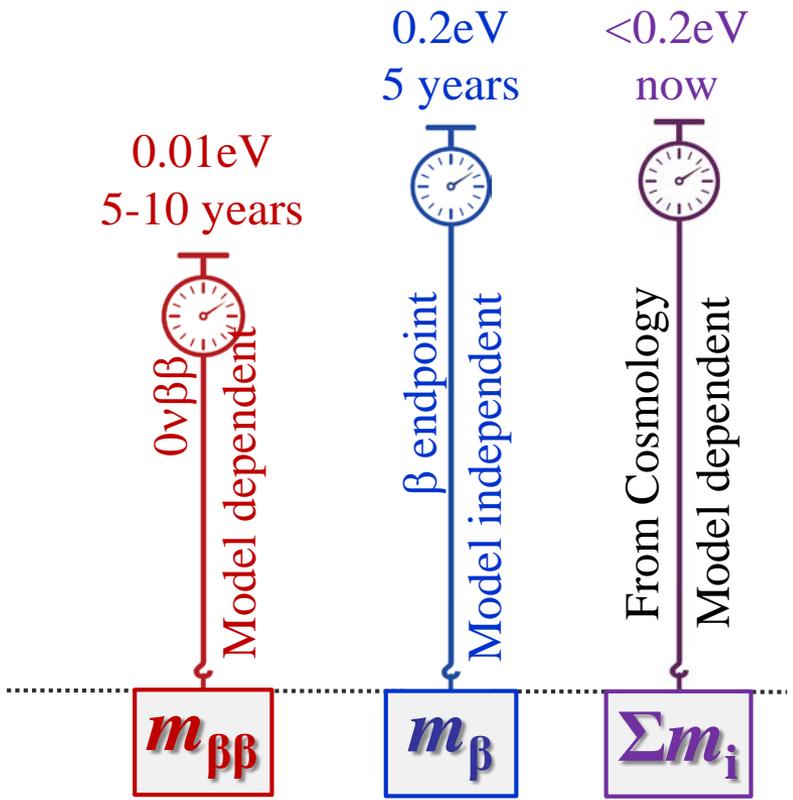
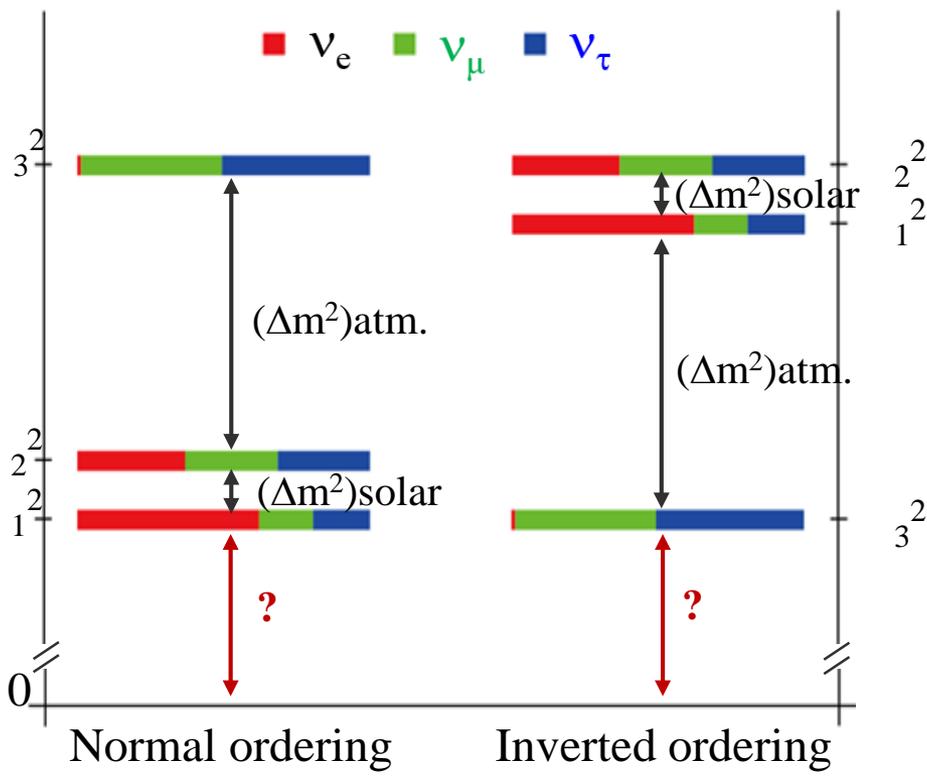
### $m$ of $\nu_e$ from $\beta$ end-point

$$m_\beta = \sqrt{\sum_{i=1}^3 |U_{ei}|^2 m_i^2}$$

### Observational cosmology

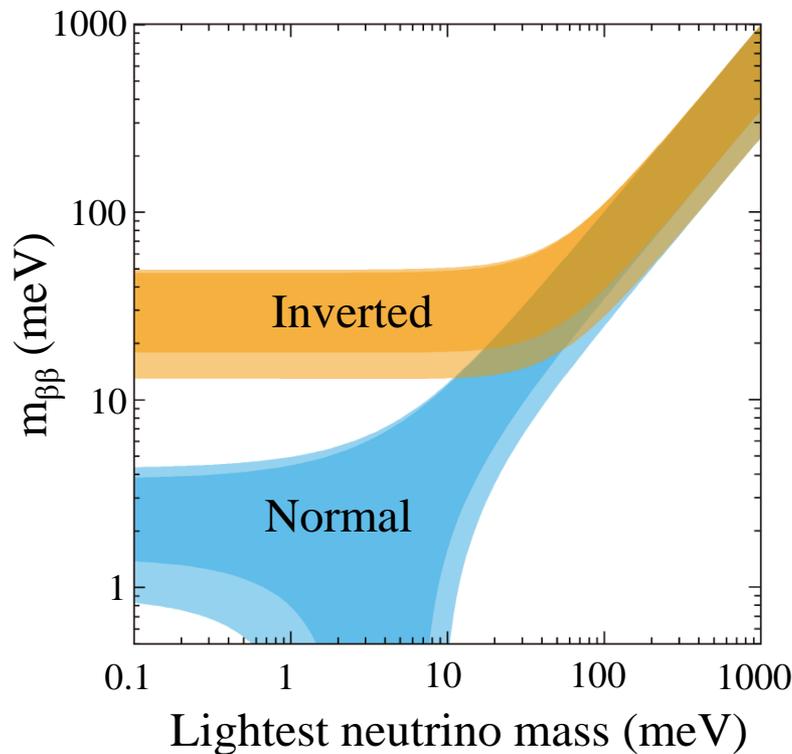
$$\sum m_\nu = \sum m_i$$

# Present knowledge of $\nu$ mass pattern & scale



# $m_{\beta\beta}$ allowed region: “usual” plot

$$m_{\beta\beta} = \left| \sum_i^2 \right| \quad \text{Parameters with known, limit and unknown values}$$



$$1/_{1/2}^{0\nu} \propto \frac{2}{m_{\beta\beta}}$$

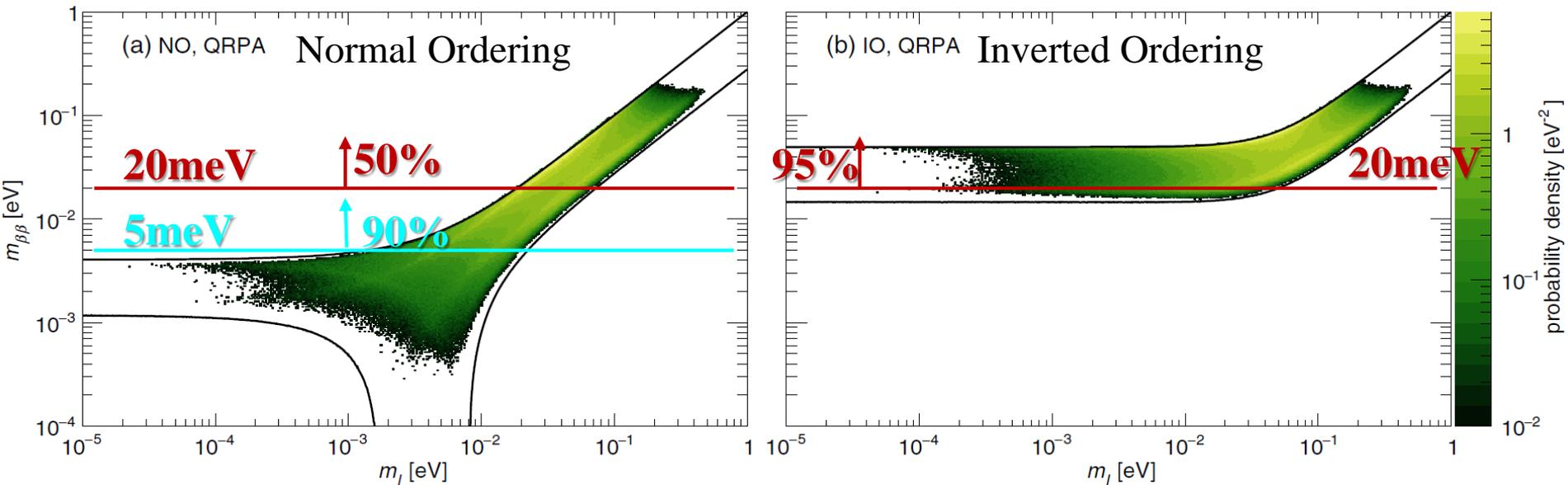
The smaller  $m_{\beta\beta}$  is the more difficult to discover  $0\nu\beta\beta$ .

- ✓ The mass ordering matters.
- ✓ The lightest  $m_i$  also matters.

# $m_{\beta\beta}$ expected region: “probable” plot

## Discovery Probability Distribution <Agostini et al PRD 96 (2017) 053001>

- ✓ Global Bayesian analysis including  $\nu$  osc.,  $m_\beta$ ,  $m_{\beta\beta}$ , and  $\Sigma m_i$
- ✓ Flat prior for the Majorana phases
- ✓ see also Caldwell, et al, PRD 96 (2017) 073001.



- ✓ **20 meV sensitivity → 50% (NO) or 95%(IO) discovery probability**  
**→ Clarification of mass ordering.**
- ✓ **5 meV sensitivity → 90% for NO**

# Physics uncertainties after $0\nu\beta\beta$ discovery

Master formula of  $(A, Z) \rightarrow (A, Z + 2) + 2e^-$

$$\Gamma_{0\nu} = \left| \sum_i \dots \right|^2 \cdot \eta^2 \quad \eta : \text{physics processes leading to Lepton number violation.}$$

Standard interpretation : Only light Majorana  $\nu$ 's lead to  $0\nu\beta\beta$

$$\Gamma_{0\nu} = \left| \sum_i \dots \right|^2 \frac{2}{\beta\beta}$$

- ✓ **Experimental  $0\nu\beta\beta$  discovery demonstrates massive Majorana particles and Lepton number violation.**
- Other mechanisms exist leading  $0\nu\beta\beta$  in the same order as light  $\nu$  exchange mechanism,
  - Heavy  $\nu$  exchange,
  - Mechanisms with RHC,
  - Majorons, etc.
- Model dependent  $M_A$  and  $M_{0\nu}$  (NME) complication

**→  $0\nu\beta\beta$  discovery from one nucleolus is not enough for full understanding.**

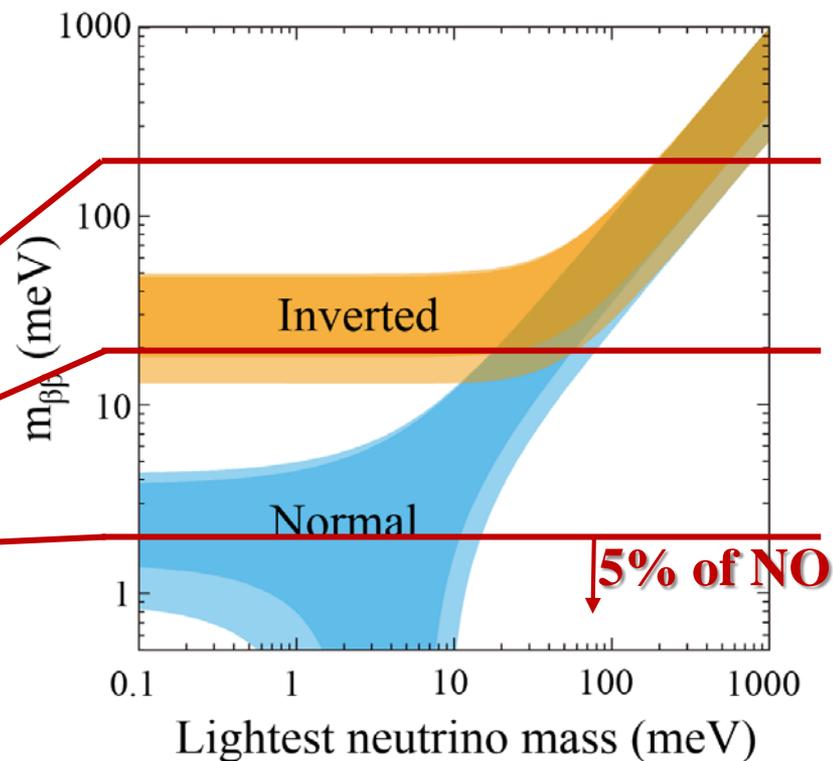
- ✓ **Full understanding requires  $0\nu\beta\beta$  results in several isotopes.**
- ✓  **$0\nu\beta\beta$  is not just a neutrino mass experiment.**

# Detection Sensitivities

# $0\nu\beta\beta$ decay rates: Simple-mined

NME dependent

Half life (years)	Decays in 10 mol (1 kg $^{100}\text{Mo}$ ) (counts/year)	$\nu$ mass scale $m_{\beta\beta}$ (meV)
$5 \times 10^{24}$	1	$\sim 200$
$5 \times 10^{26}$	0.01	$\sim 20$
$5 \times 10^{28}$	0.0001	$\sim 2$



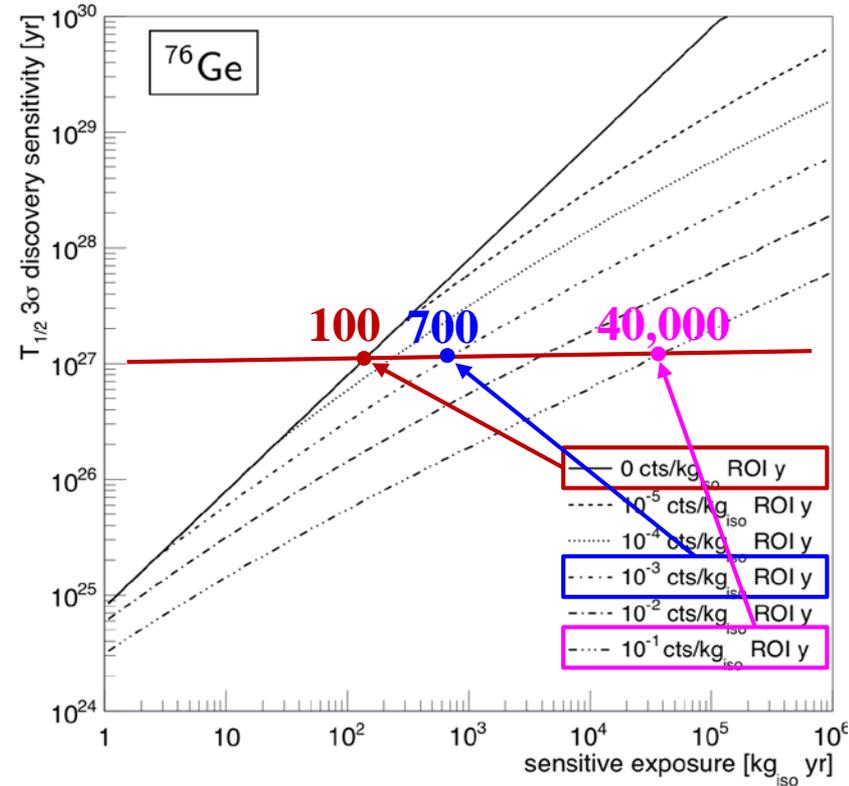
**Surely rare.**

**Matters with Bkg.,  $\Delta E$ , Exposure.**

# Experimental Sensitivity of $T_{1/2} (0\nu\beta\beta)$

For sizeable background case:

$$T_{1/2}^{0\nu}(\text{exp}) = (\ln 2) \frac{\text{Isotopic Abundance} \cdot \text{Detection Efficiency} \cdot \text{Detector Mass} \cdot \text{time}}{\text{Atomic mass} \cdot \sqrt{\text{Background level (count/keV kg year)} \cdot \text{Energy Resolution}}}$$



For “zero background” case:

(Expected background events in ROI < 1 for given time )

$$T_{1/2}^{0\nu}(\text{exp}) = (\ln 2) \frac{\text{Isotopic Abundance} \cdot \text{Detection Efficiency} \cdot \text{Detector Mass} \cdot \text{time}}{\text{Atomic mass}}$$

# Common strategies to increase sensitivity

$$\frac{0\nu}{1/2} \propto \sqrt{\frac{\cdot \text{time}}{\text{bkg} \cdot \Delta}}$$

<background case>

$$\frac{0\nu}{1/2} \propto \cdot \text{time}$$

<background-free case>

- ✓ Increase : Large detector mass, Enriched  $\beta\beta$  elements ← budget
- ✓ Increase ‘time’ : up to a few years
- ✓ Smaller  $\Delta E$  : Better energy resolution ← detector tech. LT thermal calorimeters
- ✓ Bkg. : Minimize background events in ROI
  - Underground facility (w. controls on Rn, n, dust, long-lived cosmogenics)
  - Radio-assay equipment and protocols
    - Controls on natural occurring radioactive materials (U, Th, etc. )
  - In-situ bkg. identification
    - Alphas, gammas,  $\beta\beta(2\nu)$ ,  $\mu$ - and n- induced,  $\nu$ -e scatterings
    - ← PSD, Heat/L or Charge/L detection, Veto, Shield, Topology,  $\Delta E$ ,  $\Delta t$
  - Etc. LT thermal calorimeters

Experimental approaches for  $0\nu\beta\beta$

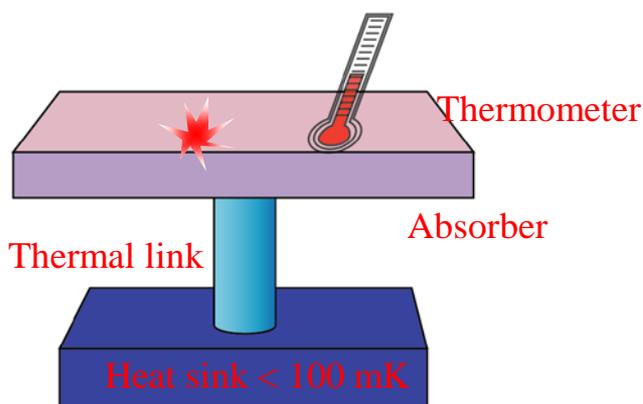
# $0\nu\beta\beta$ Experiments

Methods	Isotopes
Loaded Liquid Scintillators	$^{130}\text{Te}$ : SNO+ $^{136}\text{Xe}$ : KamLAND-Zen
Ge semiconductors	$^{76}\text{Ge}$ : GERDA Majorana Demonstrator LEGEND
TPCs (liquid, gas)	$^{136}\text{Xe}$ : EXO200, nEXO NEXT PandaX-III
Low-temperature thermal calorimeters	$^{48}\text{Ca}$ : CANDLES $^{82}\text{Se}$ : CUPID-0 $^{100}\text{Mo}$ : AMoRE, CUPID-Mo $^{130}\text{Te}$ : CUORE
Tracking chambers	$^{82}\text{Se}$ : SuperNEMO
Inorganic scintillators	$^{48}\text{Ca}$ : CANDLES

# Low Temperature Detectors

“Calorimetric measurement of heat signals at mK temperatures”

Energy absorption → Temperature



$$T - T_0 = \frac{E}{C}$$

$$\tau = \frac{C}{G}$$

Choice of thermometers for  $0\nu\beta\beta$  searches

- **Thermistors (NTD Ge)** CUORE, CUPID
- TES (Transition Edge Sensor) Light detector
- **MMC (Metallic Magnetic Calorimeter)** AMoRE
- KID (Kinetic Inductance Device) CALDER
- etc.

# Thermistors

- Doped semiconductors
  - Neutron transmuted doped (NTD) Ge thermistors
  - Ion implantation doped Si thermistors
- : 1 MΩ ~100 MΩ
- Readout: (cold) JFET
- High resolution + High linearity + Wide dynamic range + Absorber friendly
- Require very low bias current(sensitive to micro-phonics and electromagnetic interference), Slow response

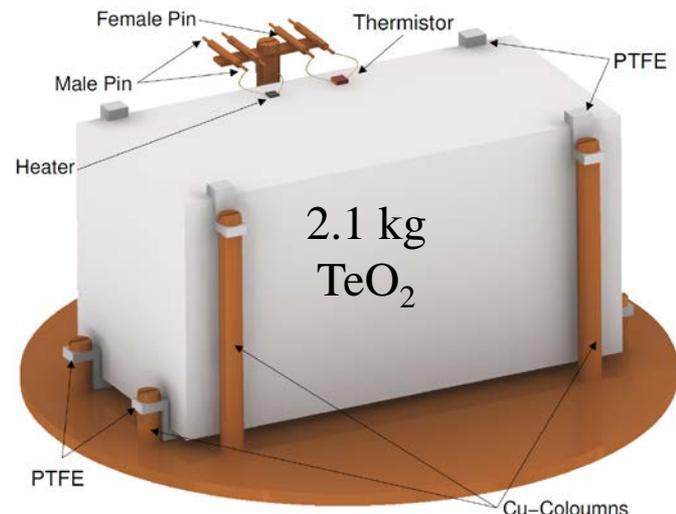
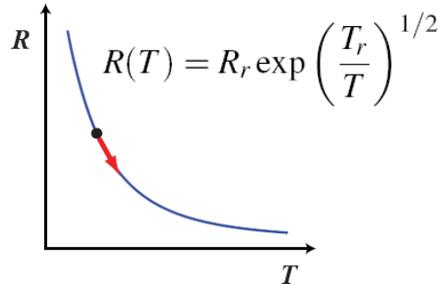
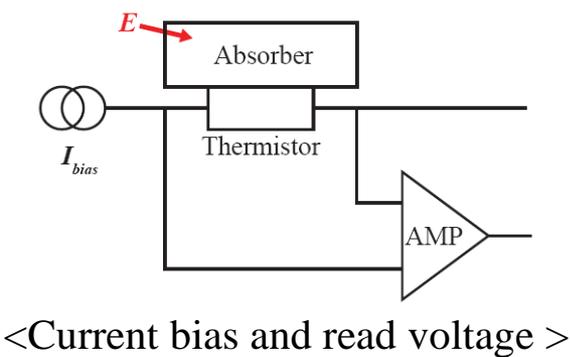


fig. from Cardani et al arXiv.1106.0568

# Metallic Magnetic Calorimeter (MMC)

- Paramagnetic alloy in a magnetic field  
Au:Er(300-1000 ppm), Ag:Er(300-1000 ppm)  
→ Magnetization variation with temperature
- Readout: SQUID
- High resolution + High linearity + Wide dynamic range + Absorber friendly + No bias heating + Relatively fast
- More wires & materials needed for SQUIDs and MMCs,

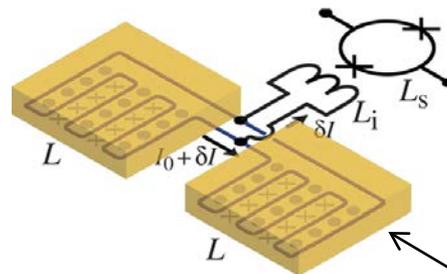
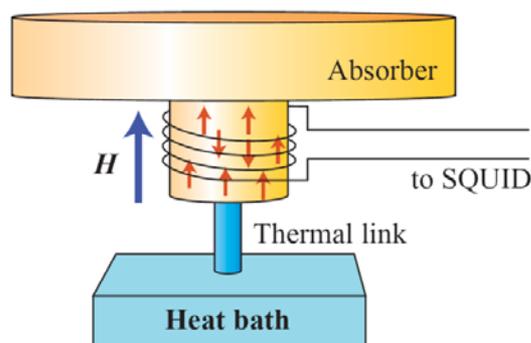
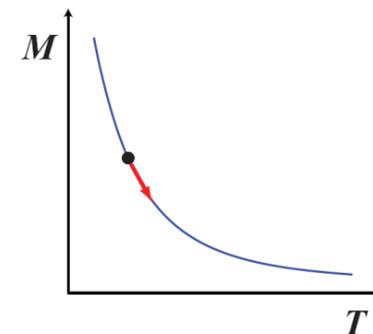
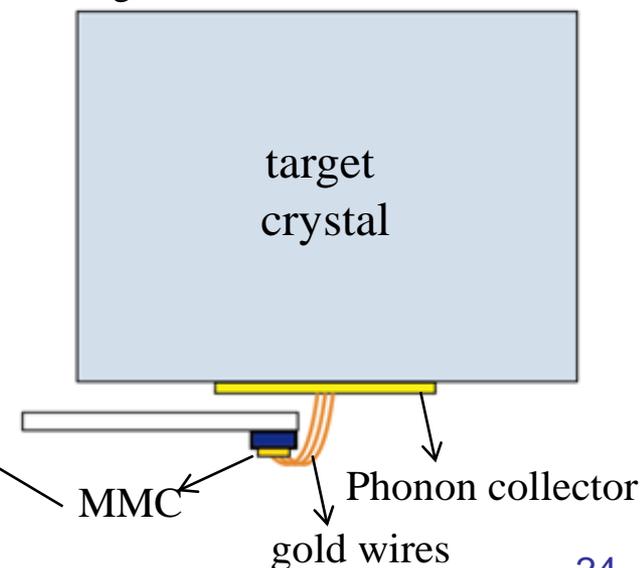
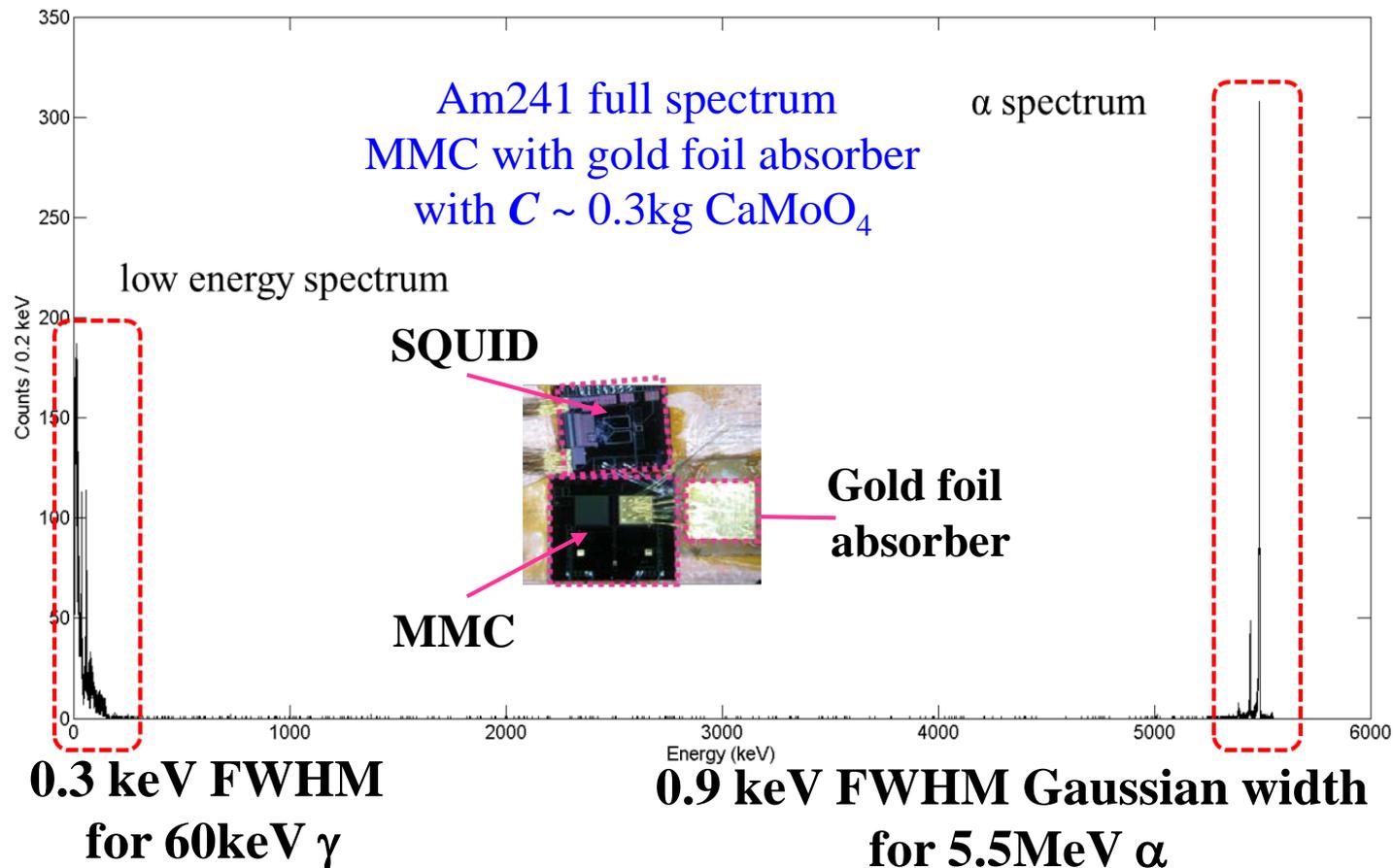


fig. from SY Oh et al SuST 2017



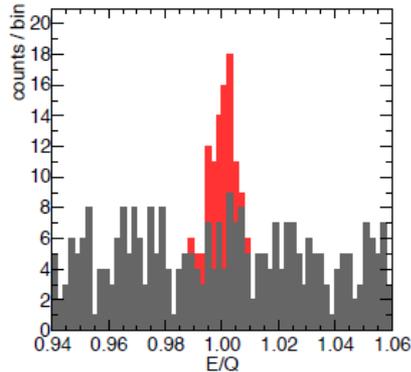
# Sensor performance

**“Superior dynamic range with high resolution”**

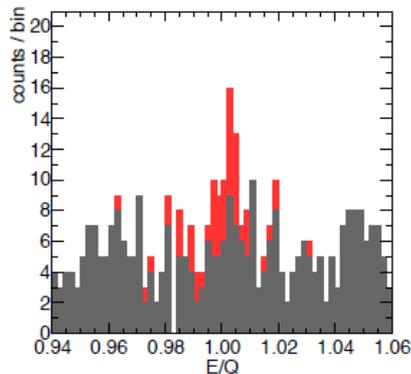


✓ A test result with an MMC.

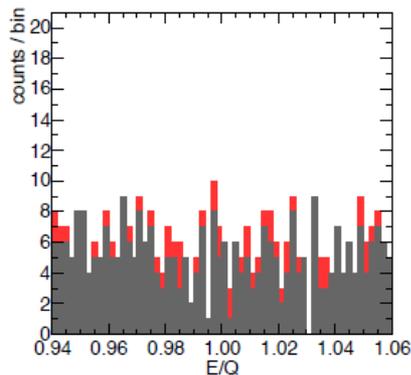
# Resolution matters.



**1 % FWHM**



**3 % FWHM**



**10 % FWHM**

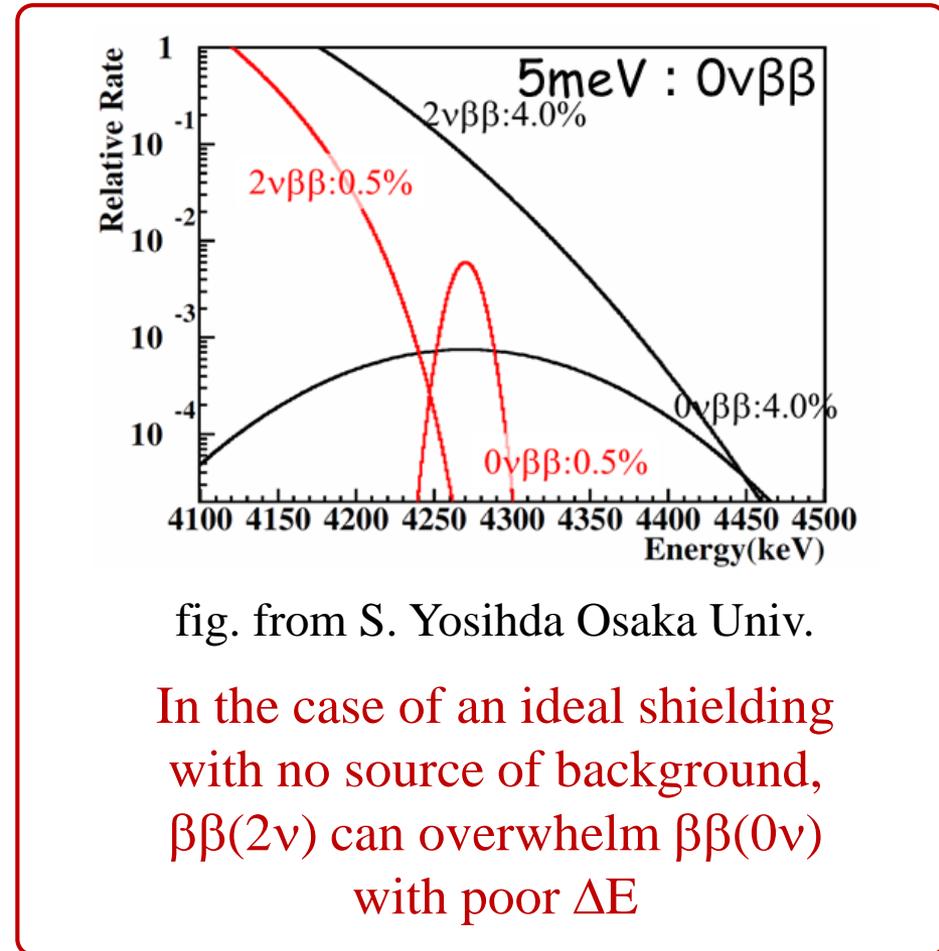


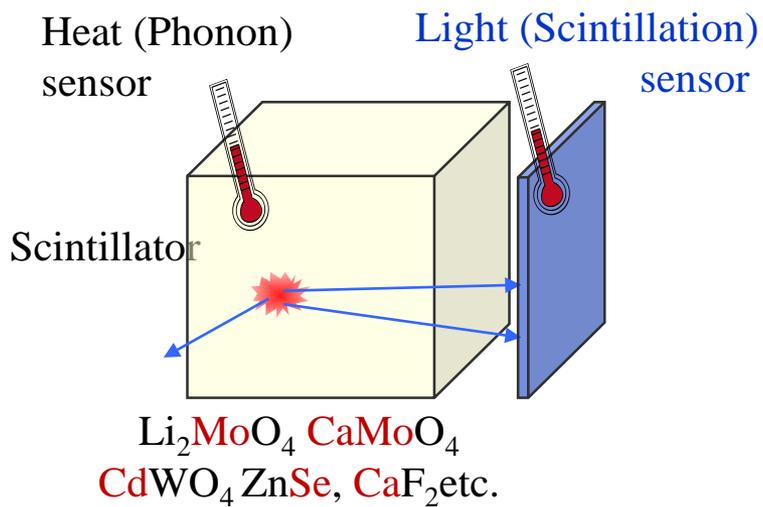
fig. from S. Yoshihda Osaka Univ.

In the case of an ideal shielding with no source of background,  $\beta\beta(2\nu)$  can overwhelm  $\beta\beta(0\nu)$  with poor  $\Delta E$

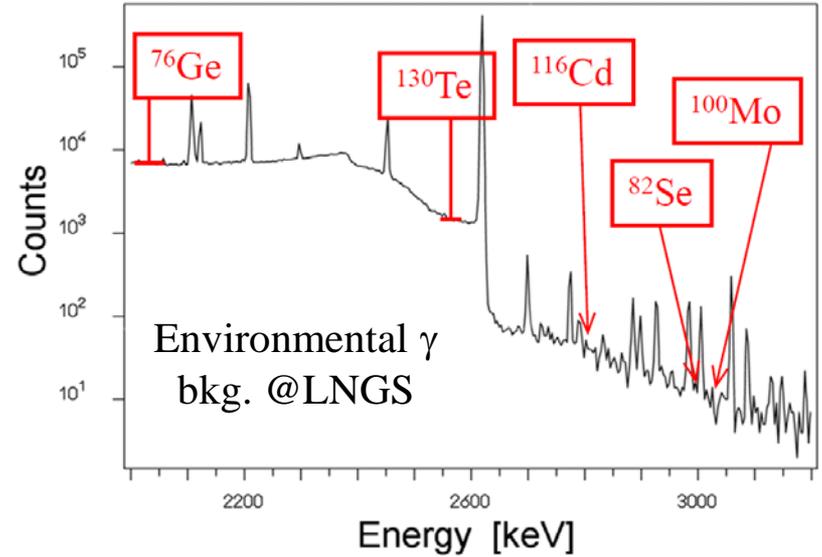
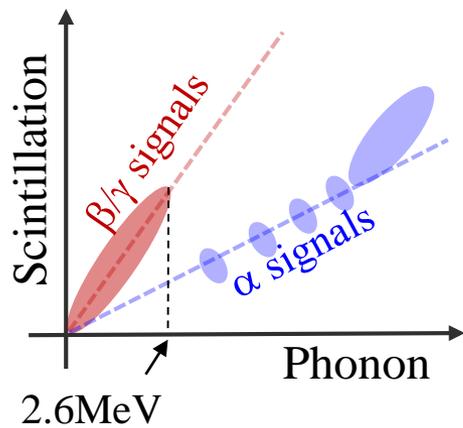
fig. from JJ Gomez-Cadenas  
XLV meeting 2017

# Simultaneous phonon-scintillation detection

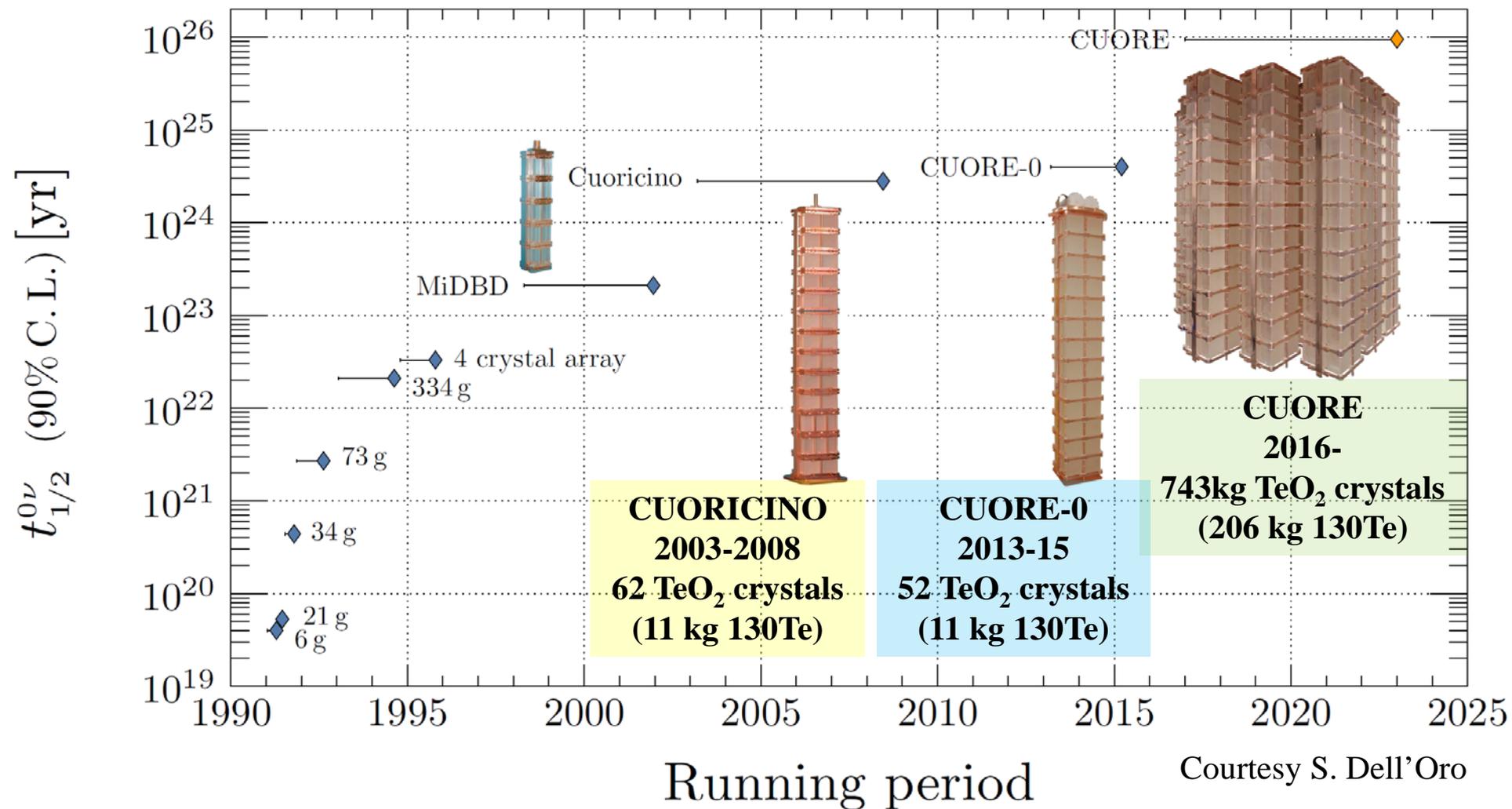
- ✓ Scintillating crystal as target material
  - ➔ Active bkg. Rejection



- ✓ Many  $\beta\beta$  nuclei test
- ✓  $Q_{\beta\beta} > 2.6 \text{ MeV}$  possible for Ca, Se, Mo
  - ➔ Low env.  $\gamma$  bkg.



# CUORE history



Courtesy S. Dell'Oro

# CUORE



- > **C**ryogenic **U**nderground **O**bservatory for **R**are **E**vents
- > Main objective:  $0\nu\beta\beta$  in  $^{130}\text{Te}$
- > 988  $\text{TeO}_2$  crystals,  $5\times 5\times 5\text{ cm}^3$  each
- > Total mass: 742 kg  $\text{TeO}_2$  (natural Te)
- >  $^{130}\text{Te}$  mass: 206 kg
- > Crystals operated as bolometers in a cryostat capable of reaching  $T < 10\text{mK}$

Present limit:

$$^{130}\text{Te } T^{0\nu} > 3.2 \times 10^{25} \text{ yr w. 90\% CI}$$

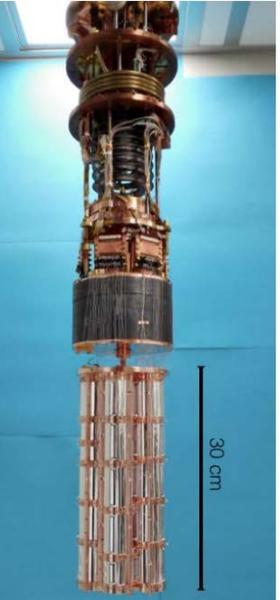
$$m_{\beta\beta} < 75\text{-}350 \text{ meV}$$

# CUPID-0

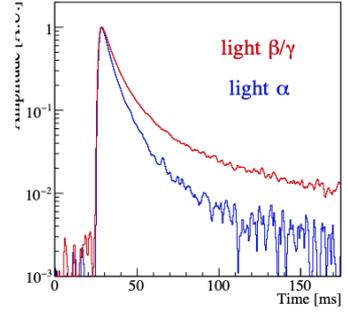
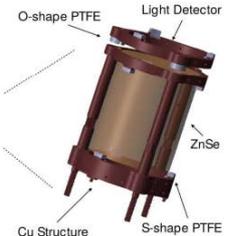
**CUPID: CUORE Upgrade with Particle Identification**

## CUPID-0: the first demonstrator

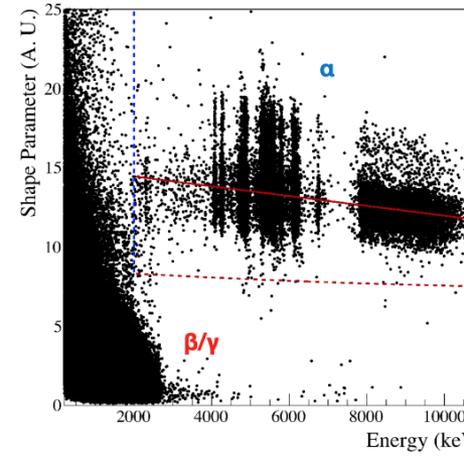
## $\alpha$ rejection with light shape



- 26 ZnSe (24 95% enriched + 2 natural)
- $^{82}\text{Se}$   $0\nu\beta\beta$  decay Q-Value: 2998 keV
- 31 Ge slabs (Light Detector)
- Arranged in 5 towers -> ( $3.8 \times 10^{25}$   $^{82}\text{Se}$  nuclei)



$\Rightarrow$  light signal shape depends on particle type

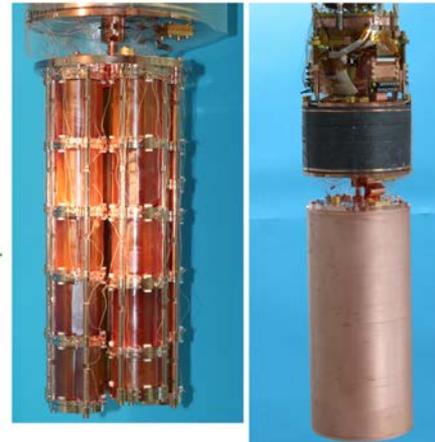
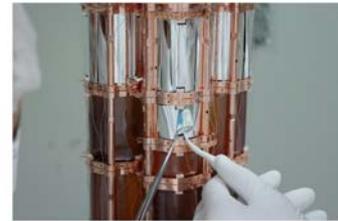


- Assembled in an underground radon free clean-room @ LNGS
- Hosted in the same CUORE-0 dilution refrigerator (Hall A)

6

$$T_{1/2}^{2\nu} = [8.60 \pm 0.03(\text{stat.}) \begin{matrix} +0.17 \\ -0.10 \end{matrix}(\text{syst.})] \times 10^{19} \text{ yr}$$

## CUPID-0 upgrade: 2019



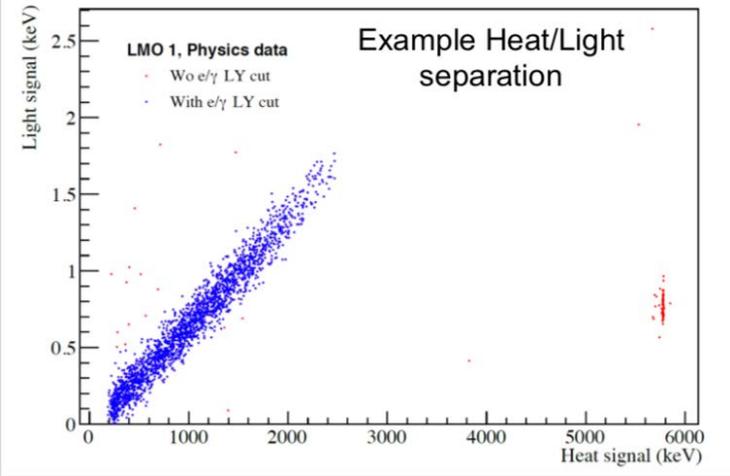
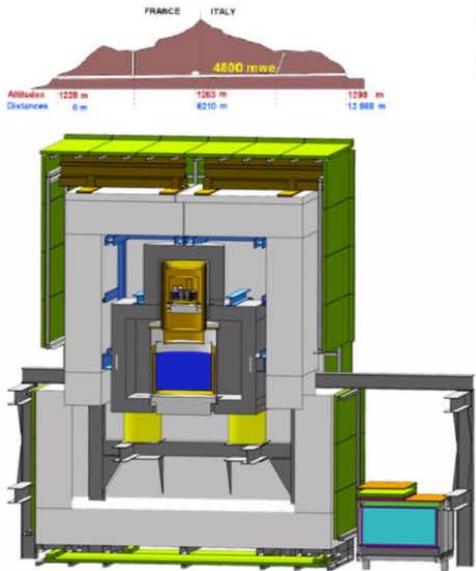
January 2019: stop data taking for a major detector upgrade:

- Remove the reflective foils
- Install a new clean copper shield
- Install a muon veto

# CUPID-Mo

## First data from the CUPID-Mo neutrinoless double beta decay experiment

Benjamin Schmidt for the CUPID-Mo Collaboration



Now: accumulated > 1 kg\*yr of physics data  
 $T_{1/2} > 3 \cdot 10^{23}$  yr at 90% C.L with  
 ~0.5 kg\*yr exposure (~0.27 kg\*yr of  $^{100}\text{Mo}$ ), 81% signal acceptance

# The AMoRE Project

AMORE: ‘Love’ in Italian



AMORE “A cosmetic company in Korea”  
Their factory is located in Daejeon.

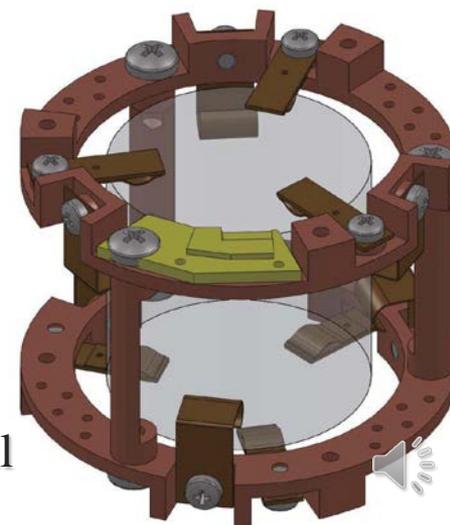


AMoRE: **A**dvanced **Mo**-based **R**are process **E**xperiment

to search for neutrinoless double decay of  $^{100}\text{Mo}$   
using **cryogenic  $\text{X}^{100}\text{MoO}_4$  detectors**

**That’s AMoRE for us!**

Ø4cmx4cm  
CaMoO<sub>4</sub> crystal



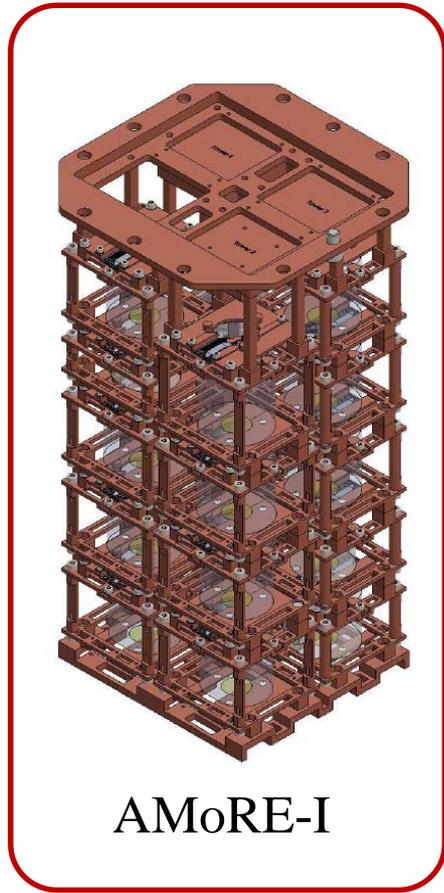
# AMoRE collaboration (since 2009)



# AMoRE Progress

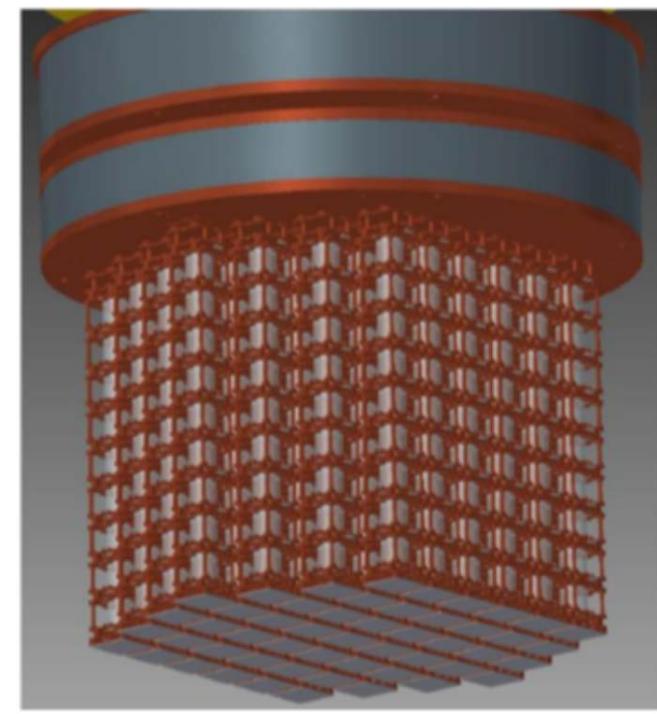


AMoRE-Pilot  
- 2018



AMoRE-I

Now

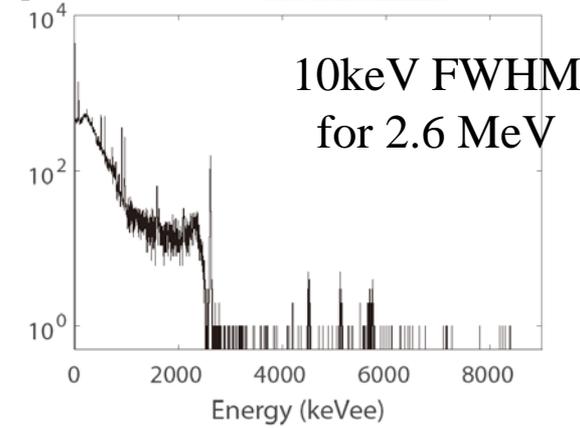
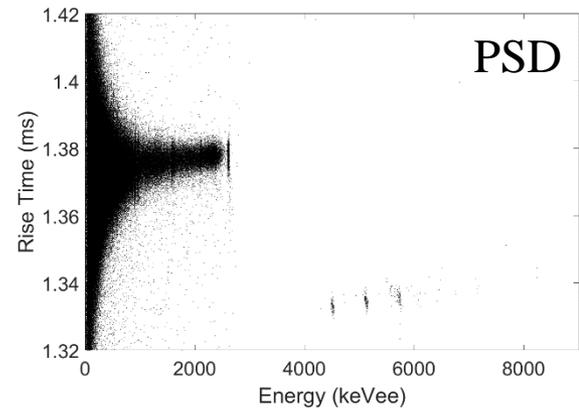
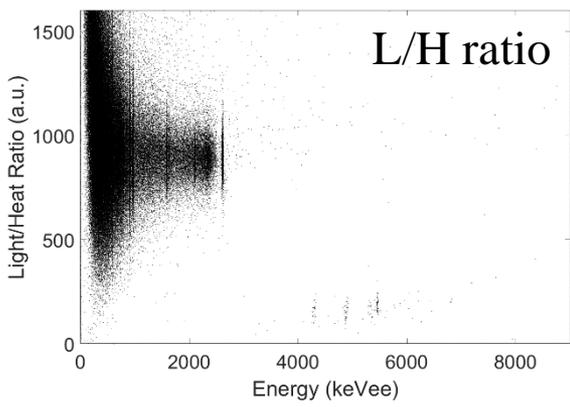
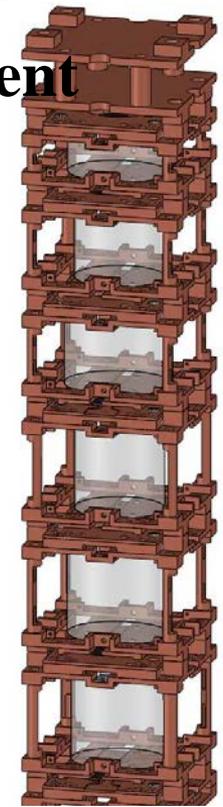
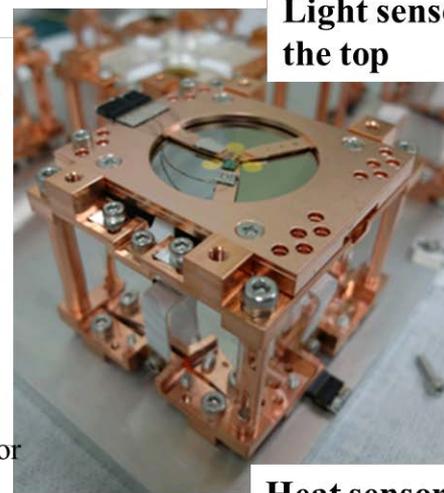
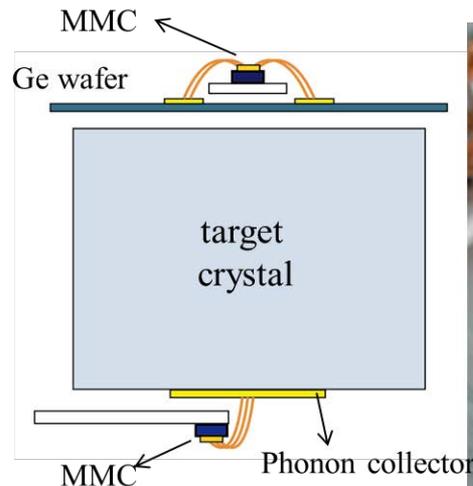
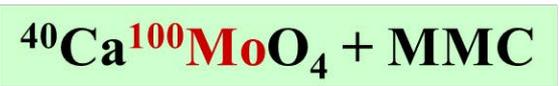


AMoRE-II  
Being prepared

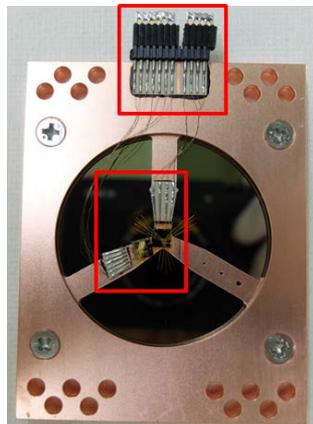
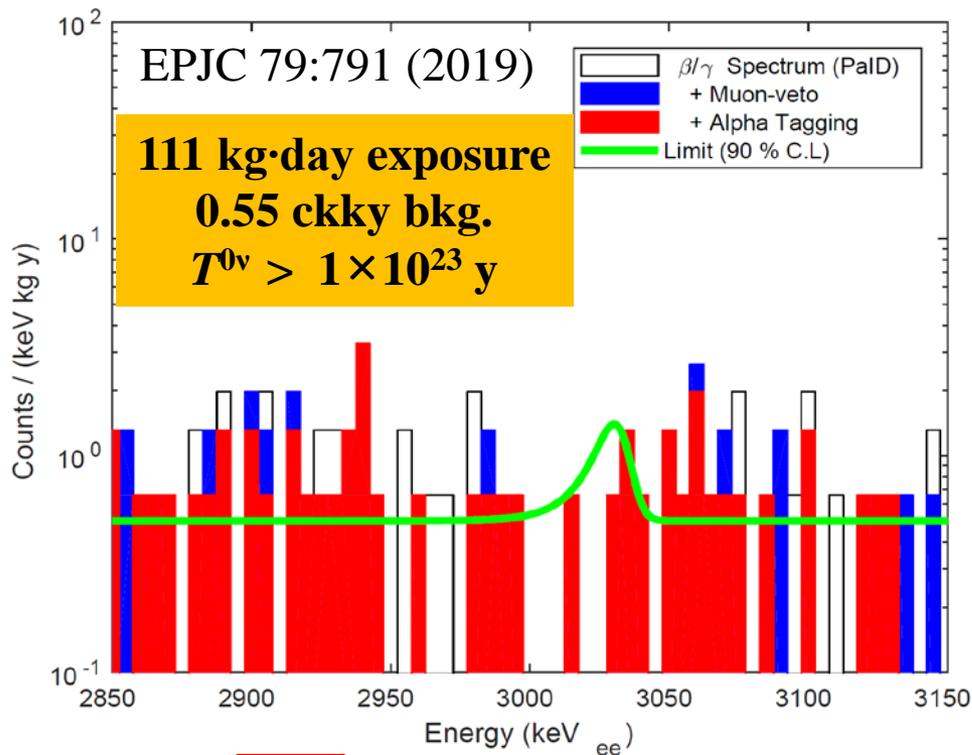
Single module

# AMoRE

AMoRE: **A**dvanced **Mo**-based **R**are process **E**xperiment



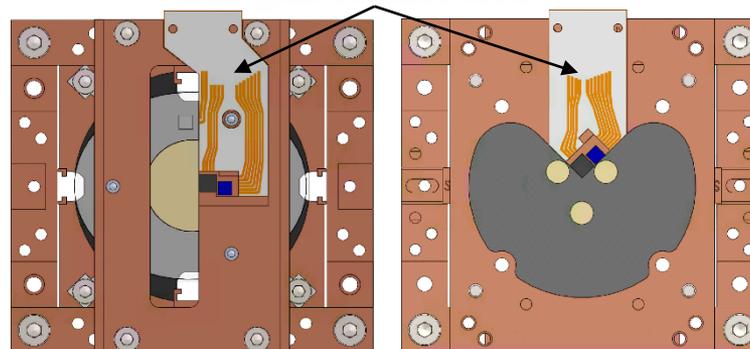
# First Pilot result, + Upgrades



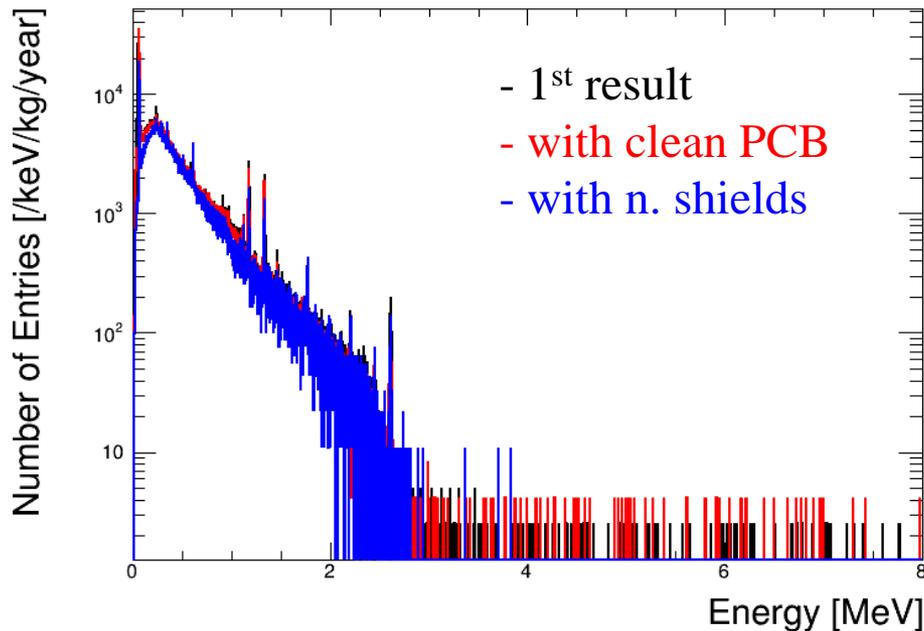
Bkg. from PCB and pin connectors

## Bkg. improvement

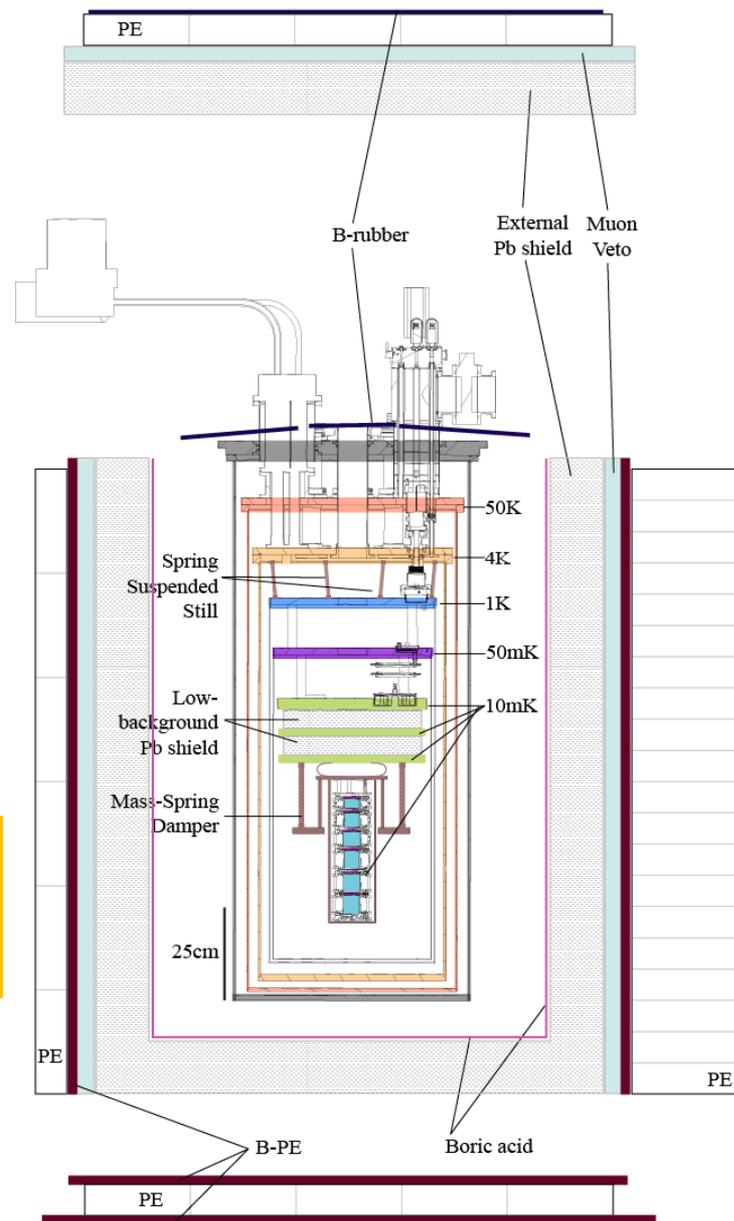
Clean PCB (tested)  
no pin connectors



# Upgrade: + Neutron shields

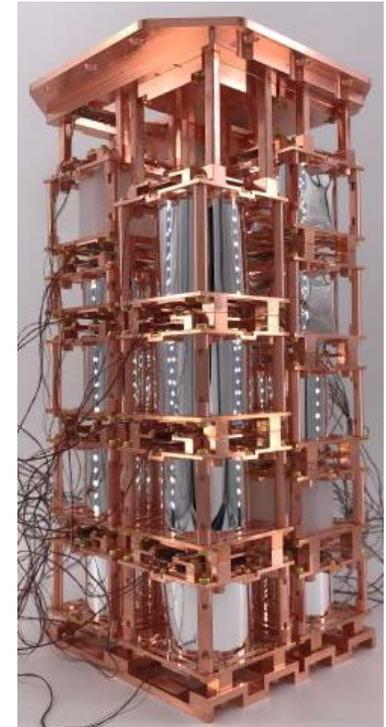
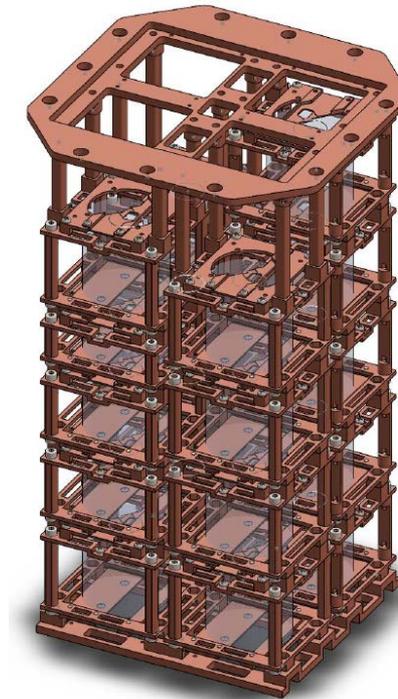


**Bkg. rate reduced by 3~9 times from the 1<sup>st</sup> result.  
New results are to be updated.**



# AMoRE-I detector

- 18 crystals
  - 36 MMC channels
  - CMOs ( $^{48\text{depl}}\text{Ca}^{100}\text{MoO}_4$ ): 13
    - 6 from Pilot (Fomos): 1.886 kg
    - 7 new crystals (Fomos): 2.696 kg
  - LMOs ( $\text{Li}_2^{100}\text{MoO}_4$ )
    - 4 (NIIC) + 1 (CUP): 1.609 kg
  - Total crystal mass 6.193 kg
  - ~ 3.1 kg of  $^{100}\text{Mo}$



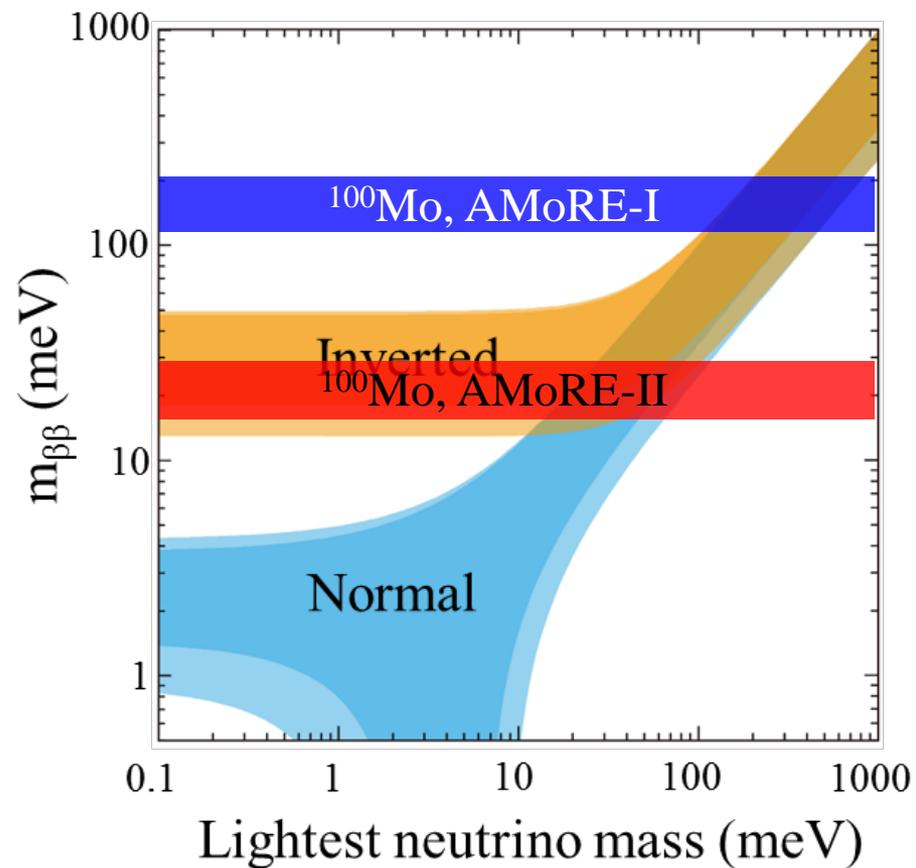
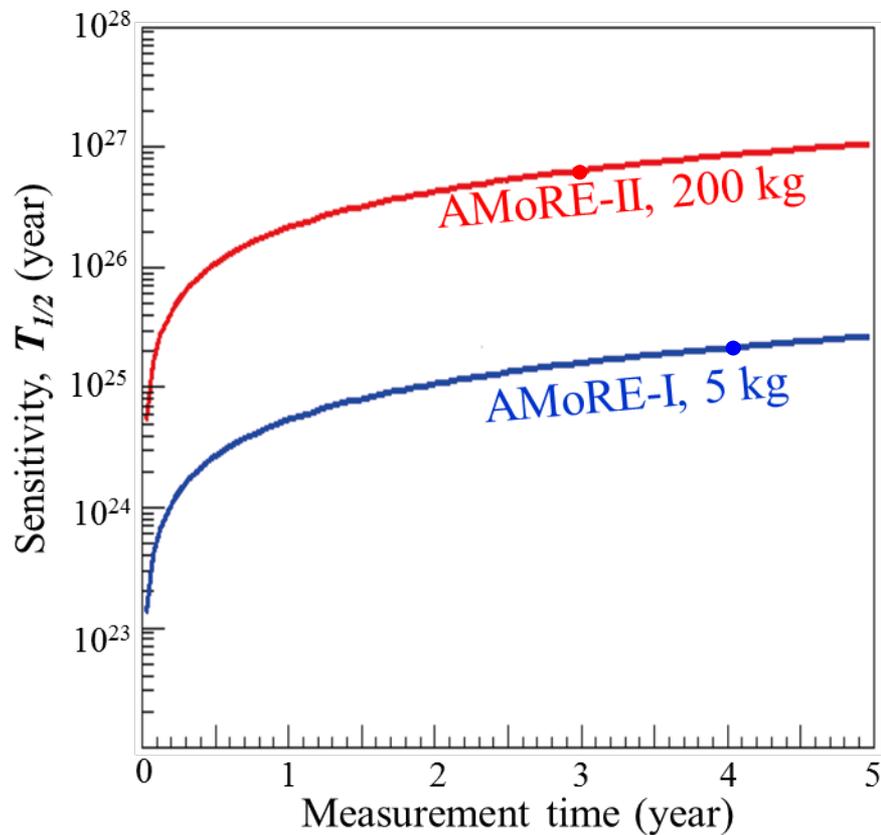
- **Science run began Dec. 2020.**

# Schedule of the AMoRE project

- MMC technology for heat and light measurement
- Crystal:  $^{40}\text{Ca}^{100}\text{MoO}_4$ , doubly enriched scintillating crystals (Pilot & I)  
For Phase II:  $\text{X}^{100}\text{MoO}_4$  (X: Li, Na, or Pb)
- Zero background in ROI
- Location: Y2L (Pilot, I) and a new deeper place (Yemilab at Jeongsun)

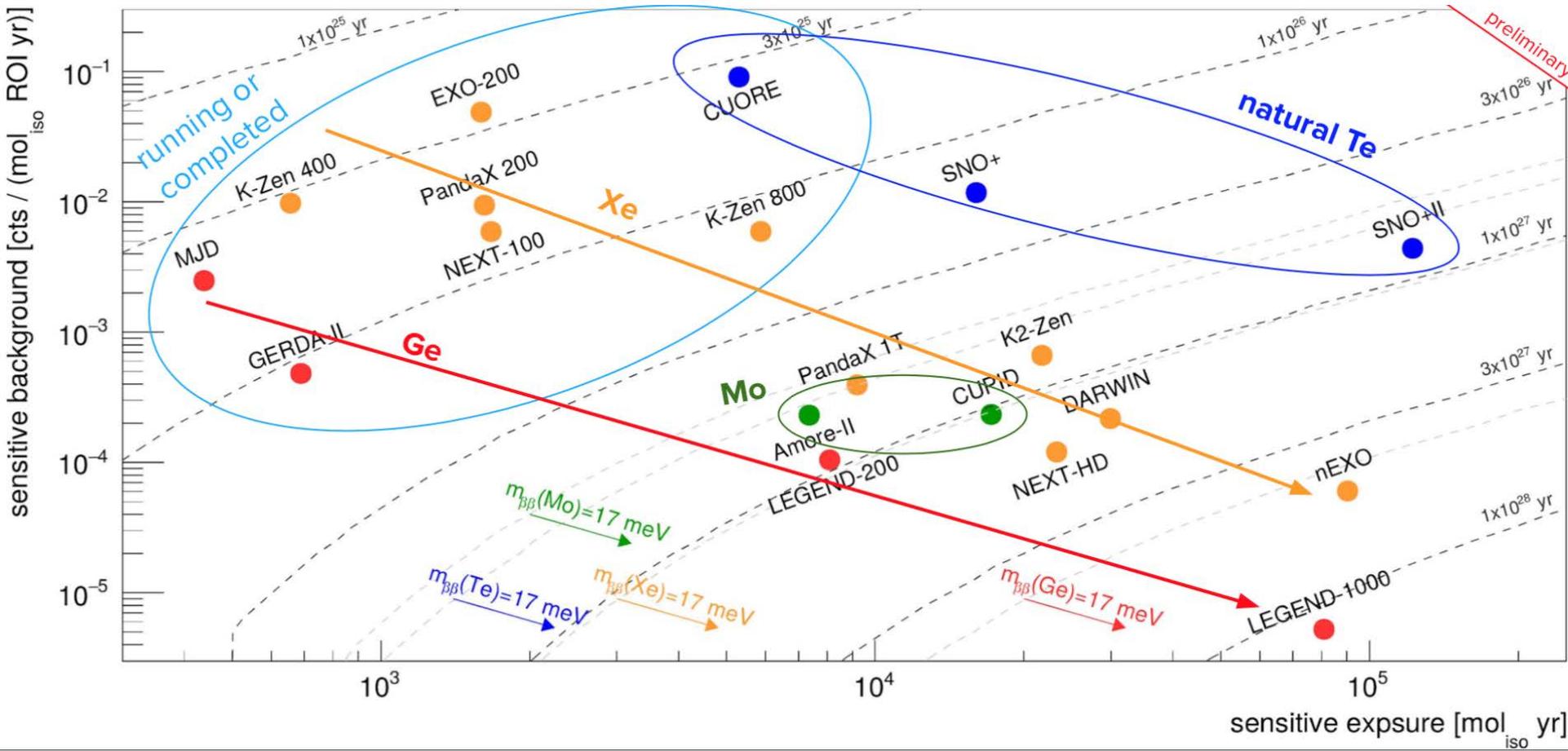
	Phase I	Phase II
Mass	~6 kg	~40 (200) kg
MMC Channel	36	1000
Required background (ckky)	0.0015	0.0001
Sensitivity( $T_{1/2}$ ) (year)	$\sim 10^{25}$	$\sim 5 \times 10^{26}$
Sensitivity( $m_{\beta\beta}$ ) (meV)	120-200	17-29
Location	Y2L	Yemilab
Schedule	2020 ~	2022 (2023) ~

# AMoRE Sensitivity



# Summary

# Current and future sensitivities



<M. Agostini 2019>

# Summary

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- $0\nu\beta\beta$  discovery is feasible in the following decade.
- NO does not mean no hope for  $0\nu\beta\beta$  discovery. (Sensitivity should reach 20meV)
- It requires many experiments with various isotopes and different detection methods.
- $0\nu\beta\beta$  search is not just a  $\nu$  mass experiment but also studies firmamental understandings of particle physics.
- Amazing progress has been realized in many projects.

- AMoRE is a promising  $0\nu\beta\beta$  experiment.
- Stay tuned to AMoRE